

**I-XE ANALYSES OF TAGISH LAKE MAGNETITE AND MONAHANS HALITE.** A. Busfield<sup>1</sup>, J. D. Gilmour<sup>1</sup>, J. A. Whitby<sup>2</sup>, S. B. Simon<sup>3</sup>, L. Grossman<sup>3</sup>, C. C. Tang<sup>4</sup> and G. Turner<sup>1</sup>, <sup>1</sup>Dept. Earth Sciences, University of Manchester, Manchester M13 9PL, UK (abusfield@fs1.ge.man.ac.uk), <sup>2</sup>Physikalisches Institut, Universität Bern, Sidlerstrasse 5, 3012 Bern, Switzerland, <sup>3</sup>Dept. of Geophysical Sciences, 5734 S. Ellis Ave., University of Chicago, IL 60637, USA, <sup>4</sup>Daresbury Laboratory, Warrington, Cheshire WA4 4AD, UK.

**Introduction:** The I-Xe dating system is particularly applicable to aqueously-formed minerals, notable examples include halite from the H-Chondrite Zag [1] and magnetite from carbonaceous chondrites [2]. Here we report data from analysis of a magnetite-rich separate from the carbonaceous chondrite Tagish Lake and of a halite grain from the ordinary chondrite Monahans (1998) (H5).

**Tagish Lake Magnetite:** 17mg of material was magnetically separated from a disaggregated sample of the Tagish Lake meteorite. Synchrotron XRD of an aliquot using 1.75Å radiation (station 2.3, Daresbury) showed that the powder was ~70% magnetite, the rest consisting of forsterite and pyrrhotite. 200µg of the powder was neutron irradiated (with chips of Shallowater enstatite) and laser step-heating I-Xe analysis performed using the RELAX mass spectrometer [3].

Results are shown in comparison to literature data from Orgueil magnetite [2] in Fig. 1A. In contrast to magnetite from Orgueil and other carbonaceous chondrites [4], there is no evidence of excess <sup>129</sup>Xe (higher than the trapped component, which is isotopically planetary). The derived iodine concentration (10ppb) and the peak I/<sup>132</sup>Xe ratios are lower than those of Orgueil magnetite, although the release pattern of iodogenic xenon with temperature appears similar. This suggests that the dramatically different result can be attributed to <sup>129</sup>I having decayed before the magnetite formed, implying a formation age at least 50 Ma later than the formation of CI magnetite.

**Monahans Halite:** A chip of Monahans halite, mass ~20 µg, was included in the same irradiation - a step release pattern is presented in Fig. 1B. All releases were dominated by iodine-derived xenon (<sup>129</sup>Xe/<sup>132</sup>Xe > 10 consistently). This sample produced a well-defined ratio of excess <sup>129</sup>Xe/I, which corresponds to an age ~4 Ma after the Shallowater enstatite standard, giving an absolute age of ~4.562 Ga (the Rb-Sr age is 4.7 ± 0.2 Ga [5]). In contrast, data from Zag halite [1], though disturbed, suggested formation before Shallowater enstatite, placing it among the earliest known minerals. Our data demonstrate that the different events that resulted in the formation of halite/sylvite in Zag and Monahans [5] were separated by ~10 Ma, and that therefore aqueous processes on the H chondrite parent body, even if intermittent,

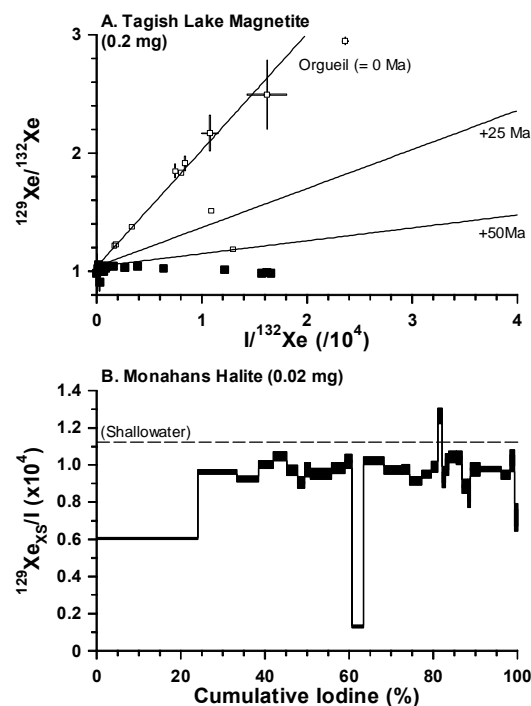


Figure 1: I-Xe data for Tagish Lake magnetite (A) and Monahans halite (B). In A, Orgueil data (open symbols - [2]) have been renormalised to <sup>132</sup>Xe. To allow comparison <sup>128</sup>Xe\* has been converted to (nominal) iodine throughout.

were ongoing for at least the first 10 Ma of solar system history.

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