

**AL-MG ISOCHRON STUDY COMBINED WITH OXYGEN ISOTOPE ANALYSIS OF THE ALLENDE TYPE B CAI, GOLFBALL.** S. Itoh<sup>1</sup>, S. B. Simon<sup>2</sup>, L. Grossman<sup>2</sup> and H. Yurimoto<sup>1</sup>. <sup>1</sup>Dept. Natural History Sciences, Hokkaido University, Sapporo, Hokkaido Japan. E-mail: sitoh@ep.sci.hokudai.ac.jp. <sup>2</sup>Dept. Geophysical Sciences, University of Chicago, Chicago, IL USA.

**Introduction:** Oxygen isotopic compositions are heterogeneously distributed within and between minerals in coarse-grained CAIs, suggesting crystallization from different generations of melt [1]. These minerals often plot on a straight line [2] or multiple lines [3] on an Al-Mg isochron diagram. A large Allende type B inclusion, Golfball, has a typical bulk composition [4] and a unique structure: a fassaite-rich mantle enclosing a melilite-rich core. From petrographic study [5], this CAI has experienced at least two melting events. Here, we report preliminary results of oxygen and Al-Mg isotopic measurements of spinel and melilite in Golfball to determine ages of the multiple melting events and the O-isotopic composition of each melt.

**Results and discussion:** Oxygen and Al-Mg isotopic analyses were performed with a Cameca ims-1270 SIMS instrument at Hokkaido University. These analytical procedures are described in detail elsewhere [6, 7]. Spinel and melilite in this CAI have oxygen isotopic compositions that plot on the CCAM line.  $\delta^{18}\text{O} = -50$  to  $-40\text{‰}$  for spinel, and the oxygen isotopic composition of melilite is distinct in each petrographic occurrence [5]:  $\delta^{18}\text{O} = -20$  to  $-10\text{‰}$  for gehlenitic core melilite,  $\delta^{18}\text{O} = -5$  to  $0\text{‰}$  for blocky melilite in the core, and  $\delta^{18}\text{O} = \sim 0\text{‰}$  for melilite in the rim. The different oxygen isotopic values may correspond to the composition of the host melt during each crystallization event.

Two isochrons are defined by the Al-Mg data. The gehlenitic core melilite and rim melilite yield initial  $^{26}\text{Al}/^{27}\text{Al}$  ratios of  $4.9 \pm 0.8$  and  $1.9 \pm 1.2 \times 10^{-5}$ , respectively. The age difference between the two isochrons is calculated to be about 1 My. The spinel plots on the gehlenitic melilite isochron, whereas the blocky melilite is on the rim melilite isochron.

From chronological and previous petrographic [5] studies, we propose the following crystallization history of Golfball. In Step 1, spinel and other minerals crystallized from an  $^{16}\text{O}$ -rich CAI melt or vapor. In step 2, after a relatively brief interval not resolved by our data, either gehlenitic melilite (but not spinel) of step 1 underwent isotopic exchange with a relatively  $^{16}\text{O}$ -poor vapor ( $\delta^{18}\text{O} = \sim -15\text{‰}$ ) or this melilite crystallized from a melt that had done so. In step 3,  $\sim 1$  My later, the CAI of step 2 was incompletely melted; and blocky melilite crystallized from the resulting melt which had become relatively  $^{16}\text{O}$ -poor ( $\delta^{18}\text{O} = \sim 0$  to  $-10\text{‰}$ ) by isotopic exchange with a vapor during melting. In step 4, after a relatively brief interval not resolved by our data, the CAI of step 3 was melted except for spinel and melilite with  $\text{Åk}_{<40}$ , and then new melilite crystallized from a melt that had become more  $^{16}\text{O}$ -poor ( $\delta^{18}\text{O} = \sim 0\text{‰}$ ) by isotopic exchange with a vapor. The melilites of Steps 3 and 4 could have crystallized in the same step if homogenization of oxygen isotopes in the melt was incomplete during crystallization.

**References:** [1] Yurimoto H. et al. 2008. *Rev. Mineral. Geochem.* 68: 141-186. [2] MacPherson G. J. et al. 1995. *Meteoritics* 30: 365-386. [3] Itoh S. et al. 2007. *Meteor. Planet. Sci.* 42: A75. [4] Simon S. B. and Grossman L. 2004. *Geochim. Cosmochim. Acta* 68: 4237-4248. [5] Simon S. B. et al. 2005. *Meteor. Planet. Sci.* 40: 461-475. [6] Itoh S. et al. 2008. *App. Surf. Sci.* 255: 1476-1478. [7] Itoh S. et al. 2007. *Met. Planet. Sci.* 42:1241-1247.