

1 **Generalized dispersion relation and critical lines for a conservative, finite-**
2 **amplitude Rossby wave in slowly varying barotropic shear flow**

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Abstract

8

9 A form of dispersion relation is proposed for a conservative, finite-amplitude,
10 barotropic Rossby wave in slowly varying parallel shear flow. The relation is
11 expressed in terms of pseudomomentum and pseudoenergy densities of the wave
12 whose exact conservation laws are known. The zonal phase speed is given by the
13 functional derivative of pseudoenergy density with respect to pseudomomentum
14 density, wherein the effects of wave-mean flow interaction and the amplitude
15 dependence of the phase speed are implicit. This theoretical prediction is compared
16 with the observed phase speed in a numerical simulation of nonlinear barotropic
17 decay on a sphere. The theory agrees well with the observed phase speed of the
18 wave except in regions where the meridional eddy momentum flux changes sign
19 and/or where the phase speed and phase tilts change abruptly. In a significant
20 departure from standard theory, multiple critical lines are identified on each flank of
21 the jet in the simulation, with significant wave amplitude through them. It is shown
22 that wave-mean flow interaction in the generalized dispersion relation renders
23 critical lines nonsingular and permits the Rossby wave to be transmitted through
24 them, even into the region where the phase speed of the wave exceeds the zonal-
25 mean zonal wind.

26 1. Introduction

27 The Rossby wave plays a central role in the large-scale circulation of the
28 Earth's atmosphere. While early linear theory illuminated the fundamental
29 properties of the atmospheric Rossby wave (Rossby¹ 1939; Haurwitz 1940; Charney
30 and Drazin 1961; Longuet-Higgins 1964, 1965; Platzman 1968; Hoskins and Karoly
31 1981; Held 1983), extensive inquiries were also directed toward the wave's
32 interaction with the mean flow (Dickinson 1969; Matsuno 1971; Andrews and
33 McIntyre 1976; Boyd 1976; McIntyre and Palmer 1983; Killworth and McIntyre
34 1985; Haynes and McIntyre 1987). It is now widely recognized that the Rossby wave
35 can drive the mean state of the atmosphere through its radiation stress, represented
36 by the generalized Eliassen-Palm flux in the quasigeostrophic limit of the
37 Transformed Eulerian Mean (TEM) formalism (Andrews and McIntyre 1976, Edmon
38 et al. 1980, Andrews et al. 1987).

39 Although the response of the mean field to the wave forcing is well described
40 by the TEM set, the corresponding theory for the finite-amplitude Rossby wave in
41 shear flow is less well developed. General description of finite-amplitude wave-
42 mean flow interaction was considered mainly in the field theory and fluid mechanics
43 literature (Sturrock 1961; Whitham 1965; Bretherton and Garrett 1968; Hayes
44 1970) and culminated in the Generalized Lagrangian Mean (GLM) formalism of
45 Andrews and McIntyre (1978ab, see also Bühler 2009). Unfortunately these theories
46 are not readily verified by (or applied to) atmospheric data because key quantities

¹ The original Rossby paper, while almost exclusively dealing with the atmospheric Rossby wave, predated the launch of the *Journal of Meteorology* (the predecessor of the *Journal of the Atmospheric Sciences*) and was published in the *Journal of Marine Research*.

47 such as pseudomomentum and pseudoenergy densities are difficult to diagnose at
 48 finite amplitude. (A notable exception is the formalism developed by Killworth and
 49 McIntyre 1985, McIntyre and Shepherd 1987, and Haynes 1988.)

50 Recently Nakamura and Zhu (2010) derived an exact conservation law for
 51 finite-amplitude pseudomomentum (wave activity) density applicable to Rossby
 52 waves and balanced eddies, extending the generalized Eliassen-Palm theorem and
 53 hence the TEM set to arbitrary eddy amplitude. For conservative barotropic flow on
 54 a sphere, this can be written as (Nakamura and Solomon 2010, Solomon and
 55 Nakamura 2012)

$$56 \quad \frac{\partial}{\partial t}(Aa \cos \phi) = \frac{\partial}{\partial \mu}(\overline{u'v'}(1 - \mu^2)), \quad (1)$$

57 where t is time, a is the radius of the sphere, $(u, v) = (\lambda a \cos \phi, a \dot{\phi})$ is the wind
 58 vector in longitude λ and latitude ϕ , $\mu \equiv \sin \phi$, and the overbar and prime denote
 59 longitudinal average and departure from it, respectively.² The angular (zonal)
 60 pseudomomentum density $Aa \cos \phi$ is defined by

$$61 \quad \begin{aligned} A(\mu, t) a \cos \phi &= \frac{1}{2\pi} (C(Q(\mu, t)) - \bar{C}(\mu, t)) \\ &= \frac{1}{2\pi} \left(\iint_{D(Q)} q a^2 d\lambda d\mu - \iint_{D(\mu)} q a^2 d\lambda d\mu \right), \end{aligned} \quad (2)$$

62 where $C(Q(\mu, t))$ is Kelvin's circulation around the wavy contour of absolute
 63 vorticity $q(\lambda, \mu, t) = Q$ that encloses the same area as the polar cap north of μ ,
 64 whereas $\bar{C}(\mu, t) = 2\pi a \cos \phi (\bar{u}(\mu, t) + \Omega a \cos \phi)$ is Kelvin's circulation around the

² Throughout the text $\cos \phi$ is used in place of $(1 - \mu^2)^{1/2}$ where it simplifies the notation.

65 zonal circle at latitude μ , and Ω is the sphere's rotation rate. The last line in (2)
66 uses Stoke's theorem and $D(Q)$ and $D(\mu)$ denote the regions delimited by the Q -
67 contour and by the latitude circle, respectively. It is readily shown (Nakamura and
68 Zhu 2010) that $A \geq 0$ and that A converges to the familiar linear pseudomomentum
69 density [Held 1985, see also (29a) below] in the small-amplitude limit. To evaluate
70 (2) one only needs instantaneous distribution of absolute vorticity q .

71 Under conservative dynamics the first term on the right hand side of (2) is a
72 constant of motion due to Kelvin's circulation theorem, so the time derivative of (2)
73 leads to

$$74 \quad \frac{\partial}{\partial t} U_{REF}(\mu) = 0, \quad U_{REF} = \bar{u}(\mu, t) + A(\mu, t). \quad (3ab)$$

75 The reference state velocity U_{REF} is the zonal velocity of the flow that would emerge
76 if the wavy q contour were 'zonalized' without changing the enclosed area or
77 Kelvin's circulation (Nakamura and Zhu 2010, Solomon and Nakamura 2012). Note
78 that U_{REF} is steady even when \bar{u} and A are not. This fact can be exploited to derive
79 conservation of pseudoenergy density $E(\mu, t)$, the difference between the zonally
80 averaged energy density and the energy density of the reference state, both
81 measured relative to the rotating sphere:³

$$82 \quad \frac{\partial}{\partial t} E = -\frac{\partial}{\partial \mu} \left(\overline{\bar{u}u' + \frac{p'}{\rho_0} + e} \right) \left(\frac{v' \cos \phi}{a} \right), \quad (4)$$

³ Unlike angular pseudomomentum density $A \cos \phi$, pseudoenergy density E does not obey rotational invariance (analogous to Galilean invariance on the beta plane), so it is important to define the reference.

83
$$E(\mu, t) = \frac{1}{2} \overline{(u' + v')^2 + v'^2} - \frac{1}{2} U_{REF}^2 = \bar{e} - U_{REF} A + \frac{A^2}{2}, \quad (5)$$

84 where $e = (u'^2 + v'^2)/2$ is eddy kinetic energy density, p' is pressure perturbation
 85 and ρ_0 is constant density. Note we used (3b) to derive the last expression in (5).

86 Both A and E vanish if the flow is zonally symmetric and the gradient of
 87 absolute vorticity is single-signed, so they are properties of eddy (Nakamura and Zhu
 88 2010). These quantities are easy to evaluate from instantaneous data and all the
 89 foregoing results are valid for arbitrary amplitude and shape of eddies, whether
 90 wavelike or turbulent. Furthermore, it is clear from (1) and (4) that the domain
 91 integrals of angular pseudomomentum and pseudoenergy are invariant in time, since
 92 the eddy fluxes on the right-hand side vanish at the poles.

93 In this article we will utilize the above conservation laws to derive
 94 generalized dispersion relation for finite-amplitude Rossby waves in barotropic
 95 shear flows. By 'generalized' we mean that the phase speed is expressed as a
 96 functional of A and E , fundamental conserved properties of the wave at arbitrary
 97 amplitude, instead of prescribed wavenumbers. This allows one to incorporate
 98 implicitly the dependence of the phase speed on wave amplitude and the effects of
 99 wave-mean flow interaction. Since both A and E are evaluable from data, the
 100 theory can be tested by numerical simulations of finite-amplitude Rossby waves.

101 The next section outlines the derivation of the generalized dispersion relation,
 102 emphasizing the assumptions made. In section 3 the theoretical prediction of phase
 103 speed is tested with a numerical simulation of a freely decaying Rossby wave in a
 104 shear flow on a rotating sphere (Held and Phillips 1987). The accuracy of the theory

105 will be demonstrated along with some limitations. Section 4 closely examines the
106 phase speed structure and critical lines in the simulation. It will be shown that the
107 geometry of critical lines differs significantly from classical theory. We will see that
108 nonlinearity in the generalized dispersion relation removes singularity from critical
109 lines and enables finite-amplitude Rossby wave to propagate through them, even
110 into a region where the phase speed is eastward with respect to the zonal mean flow.
111 The final section summarizes the results.

112 **2. Generalized dispersion relation**

113 A well-established strategy for finding the dispersion relation and
114 conservation laws for a slowly modulated, almost plane wave is to apply variational
115 principle to the *phase-averaged Lagrangian density* (Sturrock 1961, Whitham 1965,
116 1970, Bretherton and Garrett 1968, Hayes 1970, Grimshaw 1984). Normally this
117 procedure entails finding an appropriate Lagrangian, taking the average over phase,
118 and then requiring stationarity of the average action with respect to amplitude and
119 phase to obtain the dispersion relation and conservation laws, respectively.

120 However it is not always easy to find a suitable Lagrangian for a finite-amplitude
121 wave and its physical interpretation is confounded by the fact that there may be
122 more than one Lagrangian that generate the same governing equation. To date,
123 average Lagrangian densities for the Rossby wave have been found only in the small-
124 amplitude limit (Seliger and Whitham 1968, Buchwald 1972).

125 Instead of attempting to identify the exact form of Lagrangian density, in what
126 follows we *assume* that an average Lagrangian density for the Rossby wave exists as
127 a function of local wavenumbers and frequency. Then by requiring stationarity of

128 the averaged action with respect to phase, we obtain the conservation laws in terms
 129 of Lagrangian density (Whitham 1965, Bretherton and Garrett 1968). Since the
 130 derived conservation laws should correspond to (1) and (4), by matching the terms
 131 we can determine the relationships among the (unknown) Lagrangian density,
 132 (known) pseudomomentum and pseudoenergy densities A , E [(2)(5)], and the
 133 (unknown) phase speed of the Rossby wave. Finally by eliminating the Lagrangian
 134 density from these relationships we will obtain the expression for the phase speed in
 135 terms of A and E , which we call the generalized dispersion relation.

136 Barotropic flow on a sphere is described by streamfunction $\psi(\lambda, \mu, t)$ such
 137 that $\partial\psi / \partial\lambda = va \cos\phi$, $\partial\psi / \partial\mu = -ua / \cos\phi$. Let

$$138 \quad \psi(\lambda, \mu, t) = \psi' + \Delta\bar{\psi} + \psi_{REF}, \quad (6)$$

139 where

$$140 \quad \psi_{REF}(\mu) = -a \int \frac{U_{REF}}{\cos\phi} d\mu \quad (7)$$

141 is streamfunction of the reference state defined by (3) and

$$142 \quad \Delta\bar{\psi}(\mu, t) = a \int \frac{A}{\cos\phi} d\mu \quad (8)$$

143 is the change in the zonal-mean streamfunction due to eddy, where A is defined in
 144 (2). The eddy streamfunction ψ' is assumed to be nearly plane, that is, a periodic
 145 function of phase function $\theta(\lambda, \mu, t)$ such that

$$146 \quad \psi'(\theta) = \psi'(\theta + 2\pi). \quad (9)$$

147 It is understood that the average of ψ' over 2π vanishes. The local wavenumbers
 148 and frequency are defined as the gradients of phase function

149
$$k \equiv \theta_\lambda, \quad l \equiv \theta_\mu, \quad \sigma \equiv -\theta_t, \quad (10)$$

150 where the subscripts denote partial derivative. For simplicity, we only consider
 151 waves for which (k, l, σ) is independent of longitude λ (compatible with the fact that
 152 the background flow $\Delta\bar{\psi} + \psi_{REF}$ is independent of λ). Then

153
$$\begin{aligned} k_\lambda = 0, \quad l_\lambda = \theta_{\mu\lambda} = k_\mu = 0, \quad \sigma_\lambda = \theta_{t\lambda} = -k_t = 0 \\ \Rightarrow \quad k : \text{constant}, \quad l = l(\mu, t), \quad \sigma = \sigma(\mu, t). \end{aligned} \quad (11)$$

154 To ensure periodicity (9), neither the background flow $\Delta\bar{\psi} + \psi_{REF}$ nor the wave
 155 properties can change significantly over one cycle (2π) of phase. Thus, we require

156
$$l \gg \mu_b^{-1}, \quad \sigma \gg t_b^{-1} \quad (\text{except for steady state}), \quad (12)$$

157 where μ_b and t_b are the meridional and temporal scales of the background state,
 158 respectively. [A more formal two-scale analysis is spared for the sake of space—see
 159 for example Whitham (1970, 1974 ch.14).]

160 Now we assume that the barotropic equations of motion arise from
 161 stationarity of action

162
$$\delta \iiint L d\lambda d\mu dt = 0, \quad (13)$$

163 where L is the Lagrangian density of the flow. Finding the exact form of L is beyond
 164 our scope, but it is safe to assume that L depends on the first derivatives of ψ . With
 165 the waveform given by (9) and (10), the zonal average of L , denoted by \bar{L} , should
 166 depend on $(\theta_t, \theta_\lambda, \theta_\mu) = (-\sigma, k, l)$ but not on θ or λ because any derivative of ψ' is
 167 written as ψ'_θ times the gradient of θ [(9), (10)], whereas the zonal averaging
 168 eliminates the dependence on θ or λ (due to the zonal periodicity of the flow and
 169 constant k , zonal averaging has the same effect as the phase averaging). Likewise

170 \bar{L} should depend on the amplitude of ψ'_θ , $\alpha^2(\mu, t) \equiv \overline{\psi'^2_\theta}$, as well as $A / \cos\phi$ through
 171 $\Delta\bar{\psi}$ [(8)] and $U_{REF}(\mu) / \cos\phi$ through ψ_{REF} [(7)], but not explicitly on time (the time
 172 dependence enters only through modulation of $\nabla\theta$, α^2 and A). Let us then define
 173 *perturbation* Lagrangian density \mathcal{L} as

$$174 \quad \mathcal{L}(\theta_t, \theta_\lambda, \theta_\mu, \alpha^2, A, \mu) \equiv \bar{L} - L_{REF}, \quad (14)$$

175 where $L_{REF}(\mu)$ is the Lagrangian density of the reference state. Again \mathcal{L} does not
 176 depend explicitly on λ or t because of the zonal and temporal symmetries of the
 177 reference state.

178 Conservation laws and dispersion relation for the wave are derived by
 179 requiring that the average action be stationary (Whitham 1965, Bretherton and
 180 Garrett 1968)

$$181 \quad \delta \iint \mathcal{L} d\mu dt = 0 \quad (15)$$

182 with respect to independent variables of the wave, normally phase θ and amplitude.
 183 Of the list of variables in (14), both α^2 and A concern wave amplitude, one Eulerian
 184 and the other Lagrangian, and therefore they are interdependent. Furthermore, even
 185 though neither α^2 nor A depends on θ , at finite amplitude they may be related to
 186 frequency $-\theta_t$ (Sturrock 1961; Whitham 1974 ch.14). Therefore separating phase
 187 and amplitude is not trivial at finite amplitude; for this reason we only consider
 188 variation with respect to θ as an independent variable. (In linear theory, α^2 and A
 189 are proportional to each other and they are independent of frequency. Thus
 190 variation with respect to α directly yields a dispersion relation: Whitham 1965;
 191 Bretherton and Garrett 1968; Seliger and Whitham 1968.)

192 Variation with respect to θ at fixed μ leads to the following conservation
 193 laws for pseudomomentum and pseudoenergy in the form of Euler-Lagrange
 194 equations (e.g. Salmon 2013):

$$195 \quad \frac{\partial}{\partial t} \left(\theta_\lambda \frac{\partial \mathcal{L}}{\partial \theta_t} \right) + \frac{\partial}{\partial \mu} \left(\theta_\lambda \frac{\partial \mathcal{L}}{\partial \theta_\mu} \right) = 0, \quad (16)$$

$$196 \quad \frac{\partial}{\partial t} \left(\theta_t \frac{\partial \mathcal{L}}{\partial \theta_t} - \mathcal{L} \right) + \frac{\partial}{\partial \mu} \left(\theta_t \frac{\partial \mathcal{L}}{\partial \theta_\mu} \right) = 0. \quad (17)$$

197 These conservation equations arise from the fact that \mathcal{L} does not depend explicitly
 198 on λ or t (i.e., the zonal and temporal symmetries of the reference state). Equations
 199 (16) and (17) correspond to (1) and (4), respectively, for a slowly modulated, near
 200 plane Rossby wave. Matching of terms between (1) and (16) and between (4) and
 201 (17) suggests

$$202 \quad Aa \cos \phi = \theta_\lambda \frac{\partial \mathcal{L}}{\partial \theta_t} = -k \frac{\partial \mathcal{L}}{\partial \sigma} \equiv -\frac{\partial \mathcal{L}}{\partial \omega}, \quad \omega = \frac{\sigma}{k} \quad (18)$$

$$203 \quad E = \theta_t \frac{\partial \mathcal{L}}{\partial \theta_t} - \mathcal{L} = \sigma \frac{\partial \mathcal{L}}{\partial \sigma} - \mathcal{L} = \omega \frac{\partial \mathcal{L}}{\partial \omega} - \mathcal{L} = -\omega Aa \cos \phi - \mathcal{L}, \quad (19)$$

$$204 \quad -\overline{u'v'}(1-\mu^2) = \frac{kl(1-\mu^2)\alpha^2}{a^2} = \theta_\lambda \frac{\partial \mathcal{L}}{\partial \theta_\mu} = k \frac{\partial \mathcal{L}}{\partial l}, \quad (20)$$

$$205 \quad \overline{\left(\bar{u}u' + \frac{p'}{\rho_0} + e \right) \left(\frac{v' \cos \phi}{a} \right)} = \theta_t \frac{\partial \mathcal{L}}{\partial \theta_\mu} = -\sigma \frac{\partial \mathcal{L}}{\partial l} = -\omega k \frac{\partial \mathcal{L}}{\partial l}, \quad (21)$$

206 where ω is angular (zonal) phase speed of the wave. (In what follows we will prefer
 207 to work with $c \equiv \omega a$, equivalent phase speed at the equator.) From (18)

208
$$\mathcal{L} = -\int (Aa \cos \phi) d\omega = -\int (A \cos \phi) dc,^4 \quad (22)$$

209 and from (19)

210
$$\begin{aligned} E &= -\omega Aa \cos \phi - \mathcal{L} = -cA \cos \phi + \int (A \cos \phi) dc \\ &= -\int c d(A \cos \phi). \end{aligned} \quad (23)$$

211 The derivative form of (23) is

212
$$\begin{aligned} c(\mu, t) &= -\left. \frac{\partial E}{\partial (A \cos \phi)} \right|_{k, l, \mu \text{ fixed}} \\ &= \frac{U_{REF} - A}{\cos \phi} - \frac{\partial \bar{e}}{\partial (A \cos \phi)} = \frac{\bar{u}}{\cos \phi} - \frac{\partial \bar{e}}{\partial (A \cos \phi)}, \end{aligned} \quad (24)$$

213 where we used (5) to derive the second expression. In (24) the phase speed of the
 214 Rossby wave c is given by the functional derivative of pseudoenergy density E with
 215 respect to angular pseudomomentum density $A \cos \phi$. We call (23) and (24) the
 216 generalized dispersion relation, since the definitions and conservation of $A \cos \phi$ and
 217 E [(1)-(5)] do not depend on specific shape of the wave. Admittedly we assumed a
 218 rather restrictive waveform (9)-(11) to derive the dispersion relation, but the
 219 derived result (24) may be used to find the phase speed even when the exact
 220 waveform is not known (details to be outlined below). All information about the
 221 properties of the wave and the mean flow that affect the phase speed is included in
 222 E and $A \cos \phi$, quantities easily evaluable from data.

223 The generalized dispersion relation is nonlinear in two important ways. First,

⁴ An arbitrary function of k and l may be added to the right-hand side as a constant of integration, but this function is actually zero since (19) requires $\mathcal{L} \rightarrow 0$ as $E, A \rightarrow 0$.

224 the phase speed c depends on angular pseudomomentum density $A \cos \phi$ and
 225 therefore it is amplitude-dependent. This is why the right-hand side of (23) takes an
 226 integral form, and (24) a derivative form: in other words, the $A \cos \phi - E$ relation is
 227 nonlinear. Second, the dispersion relation includes the effect of wave-mean flow
 228 interaction [(3)] through the quadratic term in A [(5)]. This term is negligible at
 229 small amplitude because in that limit both \bar{e} and $A \cos \phi$ are quadratic in (the small)
 230 wave amplitude (see section 2b below). Before evaluating (24) from data, we shall
 231 discuss meridional propagation of wave amplitude and a few useful limiting cases.

232 *(2a) meridional propagation of amplitude*

233 If rewrite (1) and (4) as

$$234 \quad \frac{\partial}{\partial t}(A a \cos \phi) + \frac{1}{a} \frac{\partial}{\partial \mu}(c_A A a \cos \phi) = 0, \quad (25)$$

$$235 \quad \frac{\partial}{\partial t} E + \frac{1}{a} \frac{\partial}{\partial \mu}(c_E E) = 0, \quad (26)$$

236 where c_A and c_E are the effective meridional transport velocities of angular
 237 pseudomomentum and pseudoenergy densities, respectively, then from (18)-(21)

$$238 \quad c_A = \frac{-\overline{u'v'}(1-\mu^2)}{A \cos \phi} = \frac{kl(1-\mu^2)\alpha^2}{a^2 A \cos \phi} = -a \frac{\partial \mathcal{L} / \partial l}{\partial \mathcal{L} / \partial \sigma}, \quad (27)$$

$$239 \quad c_E = E^{-1} \overline{\left(\bar{u}u' + \frac{p'}{\rho_0} + e \right) v' \cos \phi} = -a \frac{\sigma \partial \mathcal{L} / \partial l}{\sigma (\partial \mathcal{L} / \partial \sigma) - \mathcal{L}}. \quad (28)$$

240 As we will see below, in the small-amplitude limit $c_A \approx c_E$ and they are reduced to
 241 the meridional group velocity of the Rossby wave.

242 *(2b) small-amplitude limit*

243 In the small-amplitude limit, A and E are approximated as (e.g., Held 1985;
 244 Nakamura and Solomon 2010, appendix)

$$245 \quad A \cos \phi \approx \frac{a}{2} \frac{\overline{q'^2}}{\partial \bar{q} / \partial \mu}; \quad E \approx \bar{e} - \bar{u} A \quad (29ab)$$

246 (the A^2 term in E is negligible in this limit). Since both E and $A \cos \phi$ are quadratic
 247 in small wave amplitude, $\partial E / \partial (A \cos \phi)$ in (24) may be approximated as $E / A \cos \phi$:

$$248 \quad c(\mu, t) = -\frac{E}{A \cos \phi} \approx \frac{\bar{u}}{\cos \phi} - \frac{\bar{e}}{A \cos \phi}. \quad (30)$$

249 Therefore (23) becomes

$$250 \quad E \approx -c A \cos \phi = -\omega A a \cos \phi; \quad \mathcal{L} \approx 0, \quad (31)$$

251 namely, the amplitude dependence of phase speed in (23) is negligible and the
 252 Lagrangian density vanishes. The vanishing \mathcal{L} renders $c_E = c_A$ [(28), (27)] and also

$$253 \quad 0 \approx \delta \mathcal{L} = \frac{\partial \mathcal{L}}{\partial \sigma} \delta \sigma + \frac{\partial \mathcal{L}}{\partial l} \delta l. \quad (32)$$

254 Hence with (27) and (32)

$$255 \quad c_E = c_A = -a \frac{\partial \mathcal{L} / \partial l}{\partial \mathcal{L} / \partial \sigma} \approx a \frac{\delta \sigma}{\delta l} = c_g, \quad (33)$$

256 namely c_E and c_A both approach the meridional group velocity of the Rossby wave.

257 (2c) *Steady amplitude limit*

258 When the wave amplitude is steady, the tendency terms in (16) and (17)
 259 vanish and thus the fluxes are invariant with μ . From (16) and (20)

260
$$k \frac{\partial \mathcal{L}}{\partial l} = -\overline{u'v'}(1-\mu^2) = \frac{kl(1-\mu^2)\alpha^2(\mu)}{a^2} = \text{indep. of } \mu \quad (34)^5$$

$\Rightarrow [l(1-\mu^2)]^{-1} \propto \alpha^2(\mu),$

261 and since l is invariant in time, from (10) $l_t = \theta_{\mu} = -\sigma_{\mu} = -k\omega_{\mu} = 0$. Thus c is

262 independent of latitude. The dispersion relation (24) becomes

263
$$c = \frac{\bar{u}}{\cos\phi} - \frac{\partial \bar{e}}{\partial(A\cos\phi)}, \quad (35)$$

264 where c is constant but the terms on the right-hand side are slowly varying

265 functions of μ . One may consider (34)-(35) as the finite-amplitude extension to the

266 Wentzel-Kramers-Brillouin (WKB) dispersion relation for the steady Rossby wave.

267 *(2d) Beta-plane approximation*

268 Translating the foregoing results for the beta plane is straightforward: one

269 simply needs to replace $(a\lambda, a\mu, \cos\phi, 1-\mu^2)$ with $(x, y, 1, 1)$. For example, (24)

270 becomes

271
$$c(y,t) = -\left. \frac{\partial E}{\partial A} \right|_{k,l,y \text{ fixed}} = \bar{u}(y,t) - \frac{\partial \bar{e}}{\partial A}. \quad (36)$$

272 It is well known that plane wave

273
$$\psi'(x,y,t) = B \cos(k(x-ct) + ly) \quad (37)$$

274 is an exact solution of the nonlinear barotropic vorticity equation on the infinite beta

275 plane if the background flow is uniform ($\bar{u} = u_0$) and if the phase speed c is given by

⁵ If the wave amplitude is steady globally, $\overline{u'v'}(1-\mu^2) = 0$ everywhere since the flux vanishes at the poles. Thus (34) should be applied only locally. (This is not the case for an infinite beta plane.)

276
$$c = u_0 - \frac{\beta}{k^2 + l^2}, \quad (38)$$

277 where β is the constant gradient in the Coriolis parameter. Because of the
 278 symmetry along the phase lines, it is straightforward to show that (the Cartesian
 279 version of) (2) leads to (appendix A)

280
$$A = \frac{\overline{(\nabla^2 \psi')^2}}{2\beta} = \frac{(k^2 + l^2)^2 B^2}{4\beta}, \quad (39)$$

281 i.e., exactly the same form as the linear pseudomomentum density [(29a)]. Since

282
$$\bar{e} = \frac{\overline{(\nabla \psi')^2}}{2} = \frac{(k^2 + l^2) B^2}{4} = \frac{\beta}{k^2 + l^2} A, \quad (40)$$

283 substitution in (36) recovers (38). Therefore, the well-known dispersion relation for
 284 the Rossby wave is derived from the generalized dispersion relation based on
 285 pseudomomentum and pseudoenergy densities. In particular, the ‘beta-effect’ is
 286 introduced through A [(39)].

287 *(2e) Diagnosing the phase speed of nonsteady wave from A and E*

288 The foregoing Lagrangian-based formalism is in fact applicable to a wide class
 289 of slowly modulated near-plane waves, not just Rossby waves. Results analogous to
 290 (18)-(33) have been reported for example by Sturrock (1961), Bretherton and
 291 Garrett (1968), and Grimshaw (1984). What is new here is that, since A and E for
 292 the Rossby wave (and balanced eddies) may be evaluated with data according to (2)
 293 and (5), we can test the validity of the finite-amplitude dispersion relation. A final
 294 obstacle in doing so is that (24) is written in terms of functional derivatives that
 295 cannot be evaluated explicitly because the precise functional dependence of \bar{e} on A

296 is unknown except for special cases like (40) in which the two quantities are
 297 proportional to each other.

298 To circumvent this difficulty, we rewrite (16) and (17) using (18)-(21) as

$$299 \quad \frac{\partial}{\partial t} \left(Aa(1-\mu^2)^{1/2} \right) = -\frac{\partial}{\partial \mu} \left(k \frac{\partial \mathcal{L}}{\partial l} \right), \quad (41)$$

$$300 \quad \frac{\partial E}{\partial t} = \frac{\partial}{\partial \mu} \left(\omega k \frac{\partial \mathcal{L}}{\partial l} \right). \quad (42)$$

301 Since the fluxes on the right-hand side vanish at the poles [(20), (21)], c may be
 302 evaluated as

$$303 \quad c(\mu, t) = \omega a = -\frac{\partial E^\dagger / \partial t}{\partial A^\dagger / \partial t},$$

$$304 \quad E^\dagger(\mu, t) \equiv \int_{-1}^{\mu} E(\mu', t) d\mu', \quad A^\dagger(\mu, t) \equiv \int_{-1}^{\mu} A(\mu', t) (1-\mu'^2)^{1/2} d\mu'. \quad (43)$$

305 Unlike the linear limit, the phase speed is not determined locally but involves
 306 integrals of $A \cos \phi$ and E over latitude. [In (43) the integrals are defined over
 307 $[-1, \mu]$ but they can be also defined over $[\mu, 1]$ without affecting the result.]

308 However where $\partial c / \partial \mu$ or $\overline{u'v'}$ vanishes it may be evaluated locally:

$$309 \quad c^* \equiv -\frac{\partial E(\mu, t) / \partial t}{(1-\mu^2)^{1/2} \partial A(\mu, t) / \partial t} = c + \frac{(\partial c / \partial \mu) (\overline{u'v'} (1-\mu^2))}{a(1-\mu^2)^{1/2} \partial A(\mu, t) / \partial t},$$

$$310 \quad c^* = c \quad \text{if} \quad \frac{\partial c}{\partial \mu} = 0 \quad \text{or} \quad \overline{u'v'} = 0. \quad (44)$$

311 In the next section we will use (43) to evaluate the phase speed of a numerically
 312 simulated finite-amplitude Rossby wave and compare the result with the directly
 313 measured phase speed to verify the accuracy of the theory.

314 **3. Rossby wave phase speed during nonlinear barotropic decay on a sphere**

315 The generalized dispersion relation derived above is now tested in a numerical
316 simulation of nonlinear barotropic decay on a sphere. The experimental setup is
317 identical with that of Held and Phillips (1987), HP87 hereafter, except that we use a
318 T170 spectral transform model with a 12th-order hyperdiffusion that damps the
319 shortest wave with the e-fold time of 0.3 day. A wave with zonal wavenumber $k = 6$
320 whose amplitude is initially centered at 45° N, slightly north to a midlatitude jet, is
321 allowed to evolve freely thereafter. This may be thought of as a crude model of
322 decaying synoptic eddies in the upper troposphere. The model solves barotropic
323 vorticity equation and angular pseudomomentum density $A \cos \phi$ is computed at a
324 regular time interval (6000s) using the method described in appendix B. Together
325 with the velocity output, we compute pseudoenergy density E according to (5).

326 Figure 1 samples snapshots of absolute vorticity in the Northern Hemisphere at
327 six stages of a simulation with initial wave amplitude (ζ_0 in HP87) of $4 \times 10^{-5} s^{-1}$.
328 This amplitude is chosen because the result exhibits a wealth of wave behaviors
329 before dissipation renders the flow nonconservative. The wave initially centered at
330 45° N creates large meandering of vorticity contours (Fig.1a), but subsequently the
331 meanders migrate southward with a characteristic eastward tilt with increasing
332 latitude (Figs.1bc). Eventually the wave breaks around 25° N and forms a critical
333 layer in which vorticity is quickly stirred (Figs.1d-f, see section 4 for more on critical
334 lines). By day 15 zonal wavenumber 12 emerges as a prominent structure in 20° -
335 30° N (Fig.1f). There is also poleward intrusion of filaments of low absolute vorticity,
336 although it is less wavelike because the filaments are strained by the cutoff vortices

337 that already exist in the initial condition.

338 Figures 2a and 2b show $A \cos \phi$ and E as functions of time and latitude. The
339 former is everywhere positive by definition, whereas the latter proves
340 predominantly negative because of the second term in the last expression of (5). The
341 two quantities are highly correlated and they demonstrate clearly that the initially
342 centralized wave activity separates into equatorward and poleward packets, until
343 they congregate around 24° and 55° N. At these latitudes $A \cos \phi$ and E attain
344 maximum amplitudes around day 9-10. Up to \sim day 6 the transport velocity for
345 pseudomomentum c_A is characterized by large negative (equatorward) values to the
346 south of 45° N and small positive (poleward) values to the north (Fig.2c). Then c_A
347 changes sign to positive in the midlatitude around day 8.5. This coincides with a
348 reversal of phase tilt due to rapid reconfiguration of phase lines (see Fig.11 below).
349 Such transition hints at reflection of the wave from the critical lines and/or
350 interference of multiple waves (though this is hardly visible in Figs.2a and 2b
351 because of very small wave activities in the midlatitude), at which point the
352 generalized dispersion relation becomes invalid since the theory is built on the
353 assumption of single wave. Figures 3a-c show $\bar{u} / \cos \phi$ and its departure from the
354 initial condition, as well as $U_{REF} / \cos \phi$ in the same coordinate as Fig.2. The zonal-
355 mean angular velocity $\bar{u} / \cos \phi$ is accelerated around 45° N as the wave exits the
356 source region, whereas it is decelerated where the wave packets arrive (compare
357 Fig.3b with Fig.2a). Since the region of acceleration is north to the axis of the initial
358 jet (35° N), the jet is displaced northward: by day 6, the axis moves to 42° N. This

359 shift does not occur in $U_{REF} / \cos \phi$, consistent with (3a), demonstrating that the
360 displacement of the jet is due to repartitioning of \bar{u} and A .

361 Global conservation of $A \cos \phi$ and E are examined in Fig.4. The domain
362 averages of these quantities are invariant under conservative dynamics and indeed
363 they remain constant up to \sim day 6, after which the average $A \cos \phi$ (E) decreases
364 (increases) significantly. This is largely due to dissipation of pseudomomentum
365 density by enhanced diffusive flux of vorticity (Nakamura and Zhu 2010) in the
366 critical layer. We therefore expect (43) to hold for at least first six days of the
367 simulation.

368 The phase speed in (43) is evaluated as the slope of the $A^\dagger - E^\dagger$ curve at each
369 latitude. Figure 5 shows scatter plots of E^\dagger (ordinate) versus A^\dagger (abscissa) at ten
370 different latitudes through the first seven days of simulation. These quantities are
371 computed by integrating $A \cos \phi$ and E meridionally [(43)], from the South Pole for
372 $15^\circ - 45^\circ$ N and from the North Pole for $50^\circ - 60^\circ$ N. As one might expect from Figs.2a
373 and 2b, the two quantities are tightly correlated and form a compact curve at all
374 latitudes. Each curve consists of 101 data points and they proceed in time from the
375 upper left to the lower right. The length of the curve is short at 15° and 60° N
376 because the wave amplitude is consistently small in the domain of integration
377 throughout the seven-day period (Figs. 2a and 2b). The curve length increases
378 dramatically from 15° to 35° N where the wave amplitude undergoes significant
379 growth over time. The curve shortens again from 35° to 45° N, as the wave
380 amplitude here decreases in time and this partially cancels the increasing trend in

381 the lower latitudes upon meridional integration. The slope of the curves in Fig.5
 382 [(43)] equals the phase speed $c = \omega a$. To the lowest order the curves appear close
 383 to being linear with similar slopes, but weak curvatures are also recognizable
 384 particularly during the early stage of the simulation, suggesting that the phase speed
 385 does depend on the wave amplitude.

386 The phase speed estimated from Fig.5 will be shown shortly, but for a
 387 validation purpose we need direct ‘observation’ of the phase speed, which we
 388 compute with the meridional velocity (v) based on the formula

$$389 \quad c_{obs}(\mu, t) \approx -a \frac{\overline{(\partial v / \partial t)(\partial v / \partial \lambda)}}{(\overline{\partial v / \partial \lambda})^2}, \quad (45)$$

390 where overbar denotes zonal average. Unlike spectral analysis, (45) may be
 391 evaluated instantaneously (the tendency term is evaluated from the difference
 392 between two consecutive outputs), and if v is simply translating in longitude at a
 393 uniform phase speed, i.e., $v = v(\lambda - \varphi(\mu, t))$, $\varphi_t = c_{obs}(\mu, t)$, then (45) is exact. Here v
 394 is an arbitrary smooth function and its amplitude need not be small. When multiple
 395 waves are present, that is, when ξ is also a function of λ , (45) defines a weighted
 396 average phase speed of all wave components. To demonstrate the accuracy of (45),
 397 we show in Fig.6 the phase migration of the meridional velocity v at 37° and 20° N as
 398 functions of longitude and time (Hovmöller diagram): shown on the left are the
 399 model output of v normalized by their instantaneous rms value across longitudes,
 400 and on the right are the reconstructions by way of translating a sine function with
 401 the phase speed calculated from (45). In the model output wavenumber 6 remains
 402 dominant throughout the simulation, although wavenumber 12 also emerges at 20° N

403 toward the end. At 37° N there is a kink in the phase lines around day 8.5, and at 20°
404 N the phase lines are discontinuous around day 6.5 and 7.5. These irregularities are
405 due to breakup and reconnection of phase lines, to be discussed more in detail in the
406 next section. The reconstructions reproduce the salient features of the model output,
407 smoothly connecting the gaps in the phase lines. Even at the end of the 15-day
408 period the placement of the phase remains nearly identical with the model output,
409 justifying the use of (45) as a surrogate for the phase speed. (We have also used the
410 eddy component of relative vorticity in place of ν to evaluate (45) and obtained
411 virtually identical results). Finally, we will also compare the phase speed estimated
412 from (43) with the linear theory [based on (30) with the small amplitude
413 approximations to A and E (29) and the instantaneous values for the zonal-mean
414 quantities].

415 Figure 7 shows the values of phase speed based on (43) (thick dots), (45)
416 (thin solid curve), and (30) (thin dots) as functions of time. At latitudes 40° N and
417 higher, the observed phase speed c_{obs} (solid curve) starts at around 20-25 ms^{-1} and
418 slowly increases to about 30 ms^{-1} at the end of the seven day period. At lower
419 latitudes the initial phase speed is much slower—in fact negative (westward) at 15°
420 and 20° N— but quickly increases to about 25 ms^{-1} in 3-4 days. This initial
421 latitudinal gradient in the phase speed is consistent with the development of phase
422 tilt between 15° and 40° N in absolute vorticity (Fig.1b). The large fluctuation in the
423 observed phase speed at 15° N is due to rapid breakup and reconnection of the phase
424 lines across a critical line (see section 4 and Fig.11). At this latitude the theoretical

425 estimate based on (43) matches the observed values only for a short period (day 1-3)
 426 and deviate significantly at other times, but the agreement improves as one moves
 427 higher in latitude ($20^\circ - 35^\circ$ N). Overall the theoretical estimate based on (43) agrees
 428 with the observed values better than the linear theory, particularly at 15° and 60° N
 429 and during the early stage of the simulation — a somewhat surprising result given
 430 the small wave amplitude in these circumstances. A possible reason for this is that
 431 the linear dispersion relation only quantifies changes in wave geometry, while
 432 initially the wave amplitude is also changing rapidly: only the finite-amplitude theory
 433 takes the effect of changing amplitude into account. On the other hand, at 45° and
 434 50° N the linear theory is much closer to the observation—another surprise
 435 considering the large initial wave amplitude here. At 45° N the phase speed
 436 predicted by (43) is highly variable and at 50° N it is out of the frame most of the time.
 437 A likely reason for these big discrepancies is that eddy momentum flux $\overline{u'v'}$ [or
 438 equivalently, c_A : see (27)] changes sign around these latitudes (Fig.2c). Since (43)
 439 may be rewritten as [with (41), (42), (20)]

$$\frac{c\overline{u'v'}}{\overline{u'v'}} = -\frac{\partial E^\dagger/\partial t}{\partial A^\dagger/\partial t}, \quad (46)$$

441 c is indefinite where $\overline{u'v'}$ vanishes and likely subject to large uncertainties as it
 442 changes sign. In this case the local formula (44) likely produces more accurate result.
 443 Apart from these caveats, it is encouraging that (43) is capable of predicting the
 444 observed phase speeds of finite-amplitude Rossby waves in parallel shear flows,
 445 particularly considering that the slowly varying assumption is not necessarily
 446 satisfied in the simulation. That said, we do not necessarily advocate using (43) as a

447 diagnostic method for c given that (45) is accurate and much easier to compute.
 448 True utility of the generalized dispersion relation lies in the fact that it can describe
 449 the effects of nonlinearity on wave dynamics as demonstrated in the next section.

450 **4. Phase speed structure and critical lines**

451 The dispersion relation (24) is built on several assumptions: (i) dynamics is
 452 conservative; (ii) the wave is near plane and consists of a single zonal harmonic; (iii)
 453 wave amplitude, phase speed, and the zonal-mean state all vary slowly in time and
 454 latitude compared with the phase. Given these restrictions, it is expected that the
 455 predicted phase speed becomes less accurate as one or more of these conditions are
 456 violated.

457 In linear wave theory, WKB assumption breaks down at turning latitudes ($l = 0$)
 458 and critical lines ($c = \bar{u} / \cos\phi$, CLs hereafter). At CL the linear wave theory itself
 459 breaks down due to singularity. The linear WKB dispersion relation for the Rossby
 460 wave on the beta plane reads [cf. (30)]

$$461 \quad \bar{u}(y) - c = \frac{\bar{e}}{A} = \frac{\beta_{\text{eff}}(y)}{k^2 + l^2(y)}, \quad (47)$$

462 where β_{eff} is the local meridional gradient of absolute vorticity. As long as β_{eff} is
 463 nonzero, (47) implies $l^2 \rightarrow \infty$ as the wave approaches a CL ($\bar{u} - c = 0$). This implies
 464 that both the meridional phase and group velocities vanish at the CL:

$$465 \quad c^y = \frac{k}{l}c \rightarrow 0, \quad c_s = \frac{\partial(kc)}{\partial l} = \frac{2kl\beta}{(k^2 + l^2)^2} \rightarrow 0 \quad \text{as } c \rightarrow \bar{u},^6 \quad (48)$$

466 i.e., an incident wave cannot reach the CL in finite time.

⁶ Although (47) is formally invalid at the CL, these limits still hold.

467 Given that the dispersion relation (24) is based on finite-amplitude theory,
 468 can we expect a different behavior at CLs? To gain some insight on this, we shall
 469 examine the spatio-temporal structure of the CLs observed in the simulation
 470 described in section 3. Since in our simulation both \bar{u} and c are changing in time
 471 and space, the locations of CLs can change. Figures 8a-8d show c_{obs} obtained from
 472 (45) (thick solid curve) and $\bar{u} / \cos\phi$ (thin solid curve) as functions of latitude for day
 473 0, 3, 5 and 7, respectively. Crossing of the two curves defines the location of a CL.
 474 Initially the two curves are well separated ($c_{obs} < \bar{u} / \cos\phi$ everywhere) so there is no
 475 CL (Fig.8a). However, because of rapid initial transformation of the wave the profile
 476 of c_{obs} changes quickly, and by day 3 the two curves cross on both flanks of the jet
 477 (Fig.8b). In fact, there are two crossings on each flank due to a sharp falloff of c_{obs} ,
 478 giving rise to two CLs sandwiching a region in which $c_{obs} > \bar{u} / \cos\phi$ (shaded, Fig.8b).
 479 For each shaded region, we call the CL facing the jet axis ‘the *inner CL*’ and the one
 480 facing away from it ‘the *outer CL*.’ The structure of c_{obs} remains relatively flat inside
 481 the jet from day 3 and 5, but in low latitudes it is highly transient: on day 5 two
 482 additional CLs appear to the south of the existing ones but they disappear by day 7
 483 (Figs.8cd). The two persistent CLs on the southern flank of the jet slowly migrate
 484 northward, from (13°, 19° N) on day 3 to (20°, 24° N) on day 7, whereas the two CLs on
 485 the northern flank remain nearly stationary at \sim (54°, 63° N). Although the initial
 486 formation of CLs is clearly due to the rapid change in the profile of c_{obs} rather than
 487 the modification of $\bar{u} / \cos\phi$, the magnitudes of $c_{obs} - \bar{u} / \cos\phi$ in the shaded regions
 488 remain comparable to the modification of $\bar{u} / \cos\phi$ (see Fig.3b), suggesting that for

489 the maintenance of mature CLs the wave-mean flow interaction is important. In fact,
490 the northward migration of the subtropical CLs is in step with the deceleration of the
491 mean flow in this region.

492 Figures 8e-8h show the corresponding evolution of $A/\cos\phi$ and c_A . Initially
493 the wave has no meridional tilt (Fig.1a) so $\overline{u'v'} = c_A = 0$ everywhere (Fig.8e), while
494 the peak of $A/\cos\phi$ is located at $\sim 45^\circ$ N. The latter separates into equatorward and
495 poleward moving packets, of which the former moves faster because of a greater
496 (negative) c_A (Figs.8f-8h, see also Fig.2a). These packets become increasingly more
497 focused because c_A decreases toward the front end of the packets. Although both
498 $A/\cos\phi$ and c_A tend to vanish in the far field, there remain significant wave
499 amplitudes at the CLs, particularly the inner CLs. On day 3 c_A is also clearly finite at
500 three of the four CLs. This is a significant departure from linear theory, which
501 predicts that the shaded regions are inaccessible to the Rossby wave packet. On the
502 contrary, the wave packet amplitudes are actually *greatest* in the shaded regions on
503 day 7 (Fig.8h). It appears as though CLs are no longer singular and capable of (at
504 least partially) transmitting the Rossby wave through them. The only thing that
505 appears singular is the near discontinuous variation of c_{obs} across the outer CLs
506 (Fig.8c,d).

507 As an attempt to build conceptual understanding of the CL behavior in the
508 nonlinear simulation, let us utilize the generalized dispersion relation developed
509 above with the beta-plane approximation. We also assume that the phase speed c is
510 constant. (In doing so we are tacitly removing the amplitude-dependence of phase

511 speed, so that wave-mean flow interaction is the sole nonlinear effect to be
 512 considered.) In this case the integral form of (36) is

$$513 \quad E + cA = \bar{e} + \frac{A^2}{2} - (U_{REF} - c)A = \bar{e} - \frac{A^2}{2} - (\bar{u} - c)A = 0.^7 \quad (49)$$

514 The above equation is defined for each latitude y and time t . Rearranging,

$$515 \quad \bar{u} - c = \frac{\bar{e}}{A} - \frac{A}{2} = \frac{k^2 + l^2}{2kl} c_A - \frac{A}{2}, \quad (50)$$

516 where the second equality uses $c_A = kl\alpha^2 / A$ [(27)] and $\bar{e} = (k^2 + l^2)\alpha^2 / 2$. Equation
 517 (50) replaces (47) in linear theory. From (50) at the CL ($\bar{u} = c$) we have

$$518 \quad c_A = \frac{k l A}{k^2 + l^2}. \quad (51)$$

519 Thus the transport velocity c_A at the CL becomes a function of wave amplitude.

520 Assuming l is finite, a wave with nonzero A will have nonzero c_A at a CL. This is
 521 consistent with the fact that both $A / \cos\phi$ and c_A are finite at many of the CLs in
 522 Fig.8. The implication is that under the dispersion relation (49), an incident wave
 523 packet can reach the CL in finite time and may be transmitted through it. Unlike
 524 linear theory, a CL does not require that l^2 diverge. In addition to the expression for
 525 c_A , (50) provides a constraint on the wave amplitude at the CL

$$526 \quad \bar{e} = \frac{A^2}{2}, \quad (52)$$

527 which represents a local maximum of \bar{e} with respect to A since [from the A -
 528 derivatives of (49)]

⁷ Since we fixed c , (49) and (50) should be viewed as equations for A , given c , U_{REF} , and \bar{e} , rather than the formulae for c .

529
$$\frac{\partial \bar{e}}{\partial A} = \bar{u} - c = 0,$$

530
$$\frac{\partial^2 \bar{e}}{\partial A^2} = \frac{\partial}{\partial A}(\bar{u} - c) = \frac{\partial}{\partial A}(U_{REF} - A - c) = -1 < 0. \quad (53)$$

531 Note that (52) cannot be obtained from the linear (small-amplitude) theory because
 532 the right-hand side would be two orders of magnitude smaller in wave amplitude
 533 than the left-hand side.

534 To illustrate how the wave solution varies across CLs, Fig.9 schematically
 535 depicts the left-hand side of (49) as a function of A at a fixed time for three different
 536 latitudes y_1, y_2, y_3 in decreasing order. For simplicity, we assume that U_{REF}
 537 monotonically decreases from y_1 to y_3 . Curve 1 corresponds to the highest latitude
 538 y_1 . The intersections of the curve with the abscissa provide two possible roots for A
 539 but let's say the smaller root $A = A_1$ materializes here. This is close to the linear
 540 solution [ignoring $A^2/2$ in (49)] represented by the intersection of the dash-dotted
 541 line and the abscissa. Note that the horizontal coordinate of the vertex is
 542 $A = U_{REF}(y_1) - c$ so $A_1 < U_{REF}(y_1) - c$ or equivalently $c < \bar{u}(y_1)$. Thus, at this latitude
 543 the wave travels westward relative to the zonal mean flow. At the second latitude y_2
 544 the wave amplitude is higher so the curve is shifted upward but also to the left, since
 545 $U_{REF}(y_2) < U_{REF}(y_1)$. The curve 2 touches the abscissa at the vertex, so A has a
 546 double root $A_2 = U_{REF}(y_2) - c$ or equivalently $c = \bar{u}(y_2)$. In another words $y = y_2$ is a
 547 CL. Now U_{REF} continues to decrease toward y_3 at which the left-hand side of (49) is
 548 described by curve 3. This curve again intersects the abscissa at two points (since A
 549 must be real the vertex of the curve cannot move above the abscissa), but this time

550 the larger root $A = A_3$ is chosen so that $A_3 > U_{REF}(y_3) - c$, or $c > \bar{u}(y_3)$, namely at this
551 latitude the wave travels *eastward* relative to the mean flow. This is a solution not
552 possible in the linear theory. Because of finite A , c_A remains nonzero from y_1 to y_3 ,
553 supporting the meridional transmission of pseudomomentum through the CL. A
554 wave solution is possible in the region $\bar{u} - c < 0$ because of wave-mean flow
555 interaction: finite-amplitude wave decelerates the mean flow to such an extent that
556 the wave propagates faster than the flow. Since the CL corresponds to a double root
557 of quadratic dispersion relation (49), it involves no mathematical singularity. The
558 picture that wave activity may be transmitted through CL at finite amplitude fits the
559 inner CLs (at $\sim 24^\circ$ and 53° N) in Figs.8b-d particularly well, where both A and c_A
560 are finite (Figs.8f-h) although c_A at 53° N is very small.

561 If CLs are nonsingular in (49), is there any singularity anywhere? One might
562 consider $U_{REF} - c = 0$ in (49) analogous to linear CL but it is not realizable at finite
563 amplitude because it requires $\bar{e} + A^2 / 2 = 0$. In Figs.8a-8d, $U_{REF} / \cos \phi$ indicated by
564 the dashed curve indeed stays above c_{obs} most of the time, except in the subtropics
565 on day 3 and in low latitudes on day 5 ($c_{obs} > U_{REF} / \cos \phi$ implies that the wave is
566 evanescent in latitude). Although (49) is nonsingular at $\bar{u} = c$ and $U_{REF} = c$, from
567 (50) it is clear that $c = \bar{u} + A / 2$ has the same mathematical characteristics as the
568 linear CL, including a vanishing c_A :

$$569 \quad c_A \rightarrow 0 \quad \text{as} \quad (\bar{u} + A / 2) - c = \frac{\bar{e}}{A} \rightarrow 0. \quad (54)$$

570 Note that $\bar{u} < \bar{u} + A / 2 < U_{REF}$. In Figs.8c and 8d c_{obs} starts to decrease sharply once

571 it reaches halfway between \bar{u} and U_{REF} at the flanks of the jet, and c_A nearly
572 vanishes there (on day 5 this occurs at $\sim 19^\circ$ and 63° N; Figs.8c and 8g). It suggests
573 the presence of singularity, and the sharp drop in c may be viewed as the wave's
574 attempt to avoid it. It also gives rise to another CL ($c = \bar{u}$), the outer CL, at about the
575 same latitude. Even though the CL itself is nonsingular, the sharp decrease of c
576 makes its location practically indistinguishable from the latitude of singularity (54).
577 Further transmission of wave activity through this CL may be limited given the small
578 c_A there. In the meantime, the decreasing c_A towards the outer CL causes
579 accumulation of A , which explains the large wave activities in the vicinity of CLs
580 (Figs.8g and 8h).

581 Since the CL geometry discussed above departs significantly from the
582 traditional view based on the linear theory, it is summarized schematically in Fig.10.
583 In the classical linear theory the wave solution is possible where $c < \bar{u}$ yet the
584 incident Rossby wave cannot reach the CL in finite time because a diverging l causes
585 the meridional group and phase velocities to vanish at the CL (Fig.10a). In finite-
586 amplitude theory, a nonzero A allows the group velocity to remain finite at the inner
587 CL, so the wave activity may be transmitted through it into the region in which $c > \bar{u}$
588 (Fig.10b). However, the wave soon encounters singularity as c approaches halfway
589 between \bar{u} and U_{REF} , where c decreases sharply and c_A nearly vanishes. It creates
590 the outer CL at the same location (Fig.10b), which transmits little wave activity
591 through it.

592 When the wave energy at the CL is large, streamlines develop closed contours
593 ('cat's eye') with substantial meridional width to an observer co-moving with the

594 wave, a favorable condition for wave breaking (Haynes and McIntyre 1987). This is
 595 the case with the inner CLs at $\sim 24^\circ$ and 53° N in Figs.8cd. Particularly around 24° N,
 596 where there was little wave activity initially, a rapid wave breaking and stirring of
 597 absolute vorticity occurs (Figs.1d-1f). The well-stirred region between $\sim 20^\circ$ and 30°
 598 N is identified as the nonlinear *critical layer* (or ‘surf zone’, McIntyre and Palmer
 599 1983). The inner CL is located at the center of the critical layer, whereas the outer CL
 600 is at the southern edge of the critical layer. Thus, the separation between the inner
 601 and outer CLs, d , is roughly (half) the width of the nonlinear critical layer. Figure 10
 602 suggests that d is approximately written as

$$603 \quad d \approx \frac{A_{crit}}{|\partial \bar{u} / \partial y|_{crit}} \approx \frac{(\bar{e}_{crit})^{1/2}}{|\partial \bar{u} / \partial y|_{crit}}, \quad (55)$$

604 where the subscript *crit* indicates characteristic values in the critical layer, and we
 605 used (52) to derive the second expression (constant factors are dropped). Thus,
 606 more energetic wave and/or weaker horizontal shear make the critical layer wider.
 607 Killworth and McIntyre (1985) show that the upper bound of d is $\approx \beta^{-1} |\partial \bar{u} / \partial y|_{crit}$.

608 Once the wave reaches the outer CL, it encounters an abrupt drop in c . Since
 609 this violates the assumption of slow variation, the dispersion relation breaks down
 610 there (even if singularity is averted). This occurs relatively early in the simulation at
 611 low latitudes; this is why the agreement between theory and observation in Fig.7 is
 612 short-lived at 15° and 20° N. Rapid change in c implies rapid change in l , or
 613 equivalently, the tilt of the wave. Figure 11 shows the phase structure (longitude-
 614 latitude) of v -velocity between day 7.8 and 8.8 of the simulation. There are visually
 615 discernible discontinuities in the meridional tilt at $\sim 15^\circ$, 20° , 35° , and 62° N. These

616 all coincide with gaps in c (not shown), and except at 35° N they coincide with the
617 outer CLs. Since discontinuity in l means discontinuity in refractive index, it implies
618 that the incident wave is partially reflected and partially refracted by the CLs. (Wave
619 reflection is something that the WKB-like solution cannot describe.) Thus, in Figs.8c
620 and 8d, the region between the two outer CLs at 20° and 62° N is shaping up to be a
621 waveguide, hosting multiply reflected waves.

622 Figure 11 shows not only discontinuous tilt but also break-up and
623 reconnection of phase lines at the outer CLs. Since the wave is traveling at distinct
624 phase speeds across the CL, the phase lines inevitably deform and eventually break,
625 and then reconnect with the next phase lines on the other side of the CL. The jump in
626 phase lines is evident in $15^\circ - 20^\circ$ N (see also Fig.4 of Haynes and McIntyre 1987) and
627 also near 35° N. The phase jump occurs rapidly and creates anomalous transient
628 behaviors in phase speed as we observed in Fig.7 at 15° N and in Fig.6 at 20° N (phase
629 discontinuities). The phase jump at 35° N around day 8.5 in Fig.11 is preceded by a
630 gap in c (Fig.8d) and it causes a reversal of the meridional tilt, hence $\overline{u'v'}$ and c_A , as
631 we have seen in Fig.2c. Apparently this midlatitude jump is caused by interference of
632 two Rossby modes in the waveguide traveling at distinct speeds due to different
633 meridional structures. Clearly, description of these rapid phase jump behaviors is
634 beyond the ability of the dispersion relation [(24)].

635 5. Summary

636 We have extended the dispersion relation for the Rossby wave in barotropic
637 shear flow to finite-amplitude regime using the exact conservation laws for

638 pseudomomentum and pseudoenergy densities (Nakamura and Zhu 2010) and the
639 well-known method based on the phase-averaged Lagrangian density for slowly
640 modulated, near plane waves (Whitham 1965, Bretherton and Garrett 1968). The
641 phase speed is expressed as a functional derivative of pseudoenergy density with
642 respect to pseudomomentum density. Despite obvious limitations (conservative
643 dynamics, single zonal harmonic, slow modulation, near plane waveform, etc.),
644 writing the dispersion relation in terms of pseudomomentum and pseudoenergy
645 densities enhances the versatility of theory since their conservation does not depend
646 on specific wave geometry or amplitude, and the effects of wave-mean flow
647 interaction and the amplitude dependence of the phase speed may be incorporated
648 seamlessly. Furthermore, since pseudomomentum and pseudoenergy densities are
649 readily evaluable, the theory is testable with numerical simulations.

650 In addition to the dispersion relation, we have developed a method to
651 estimate the phase speed of the wave directly from instantaneous data [(45)]. Data
652 sampled from the numerical simulation of nonlinear barotropic decay on a sphere
653 (HP87) with an initial zonal wavenumber 6 demonstrate that the finite-amplitude
654 theory reproduces the observed phase speed better than the linear theory as long as
655 one is away from regions in which meridional eddy momentum flux changes signs
656 and/or the phase speed is discontinuous. It is also found that the linear theory does
657 not necessarily perform better at smaller amplitude when the wave amplitude is
658 growing quickly. Consistent with previous studies (e.g., Randel and Held 1991) CLs
659 are generally located on the flanks of the jet, but their geometry differs significantly
660 from the standard linear theory: there are multiple CLs on each side of the jet axis. It

661 is shown using the generalized dispersion relation that nonlinearity removes
662 singularity from CLs and the meridional group velocity at the CL remains finite when
663 the wave amplitude is finite. As a result, the Rossby wave may be transmitted
664 through the CLs even into the region where $c > \bar{u}$.

665 In future studies we will extend the analysis to atmospheric data and examine
666 the CL geometry in the upper troposphere and its relationship to wave breaking with
667 the aid of the generalized dispersion relation.

668 **Acknowledgments**

669 This work has been supported by NSF Grant AGS-1151790.

670 **Appendix A Derivation of (39) from (37) and (2)**

671 With the plane waveform (37) and the beta-plane approximation, absolute
672 vorticity q is

$$673 \quad \begin{aligned} q(\theta, y) &= f_0 + \beta y + q'(\theta), \\ q'(\theta) &= -(k^2 + l^2)B \cos \theta, \quad \theta = k(x - ct) + ly. \end{aligned} \quad (\text{A1})$$

674 Pseudomomentum density A is defined by the beta-plane version of (2)

$$675 \quad \begin{aligned} A(y) &= \frac{1}{2\pi} \left(\iint_{q \geq Q} q(\theta, y') dy' d\theta - \iint_{y' \geq y} q(\theta, y') dy' d\theta \right) \\ &= - \overline{\int_y^{y+\xi(\theta, y)} q dy'} = - \left[q(\theta, y) \xi(\theta, y) + \frac{1}{2} \beta \xi^2(\theta, y) \right], \end{aligned} \quad (\text{A2})$$

676 where overbar denotes phase average (= zonal average) and $\xi(\theta, y)$ is the meridional
677 displacement of the contour $q = Q$ relative to y along phase line (fixed θ). Because
678 the meridional gradient of q along the phase line is constant ($= \beta$),

$$679 \quad q(\theta, y) = Q(y) - \beta \xi(\theta, y). \quad (\text{A3})$$

680 Requiring $\overline{\xi} = 0$ (displacement is area-preserving), one sees from (A3)

$$681 \quad Q(y) = f_0 + \beta y, \quad \xi(\theta) = -q'(\theta)/\beta. \quad (\text{A4})$$

682 Substituting in (A2)

$$683 \quad A = - \overline{\left(-\frac{q'^2}{\beta} + \frac{1}{2} \frac{q'^2}{\beta} \right)} = \frac{1}{2} \frac{\overline{q'^2}}{\beta} = \frac{(k^2 + l^2)^2 B^2}{4\beta}, \quad (\text{A5})$$

684 which is (39). Notice that, although (A5) is identical with what the linear theory
685 predicts, the wave amplitude B here is not assumed small.

686 **Appendix B Calculation of A**

687 One may calculate A by evaluating the area integrals in (2) with weighted box
 688 counting method. However, this method is prone to errors when the wave amplitude
 689 is small, since in that case A becomes a small difference between two large integrals.
 690 Here we use an alternative method to calculate A .

691 By taking the derivative of (2) with respect to μ twice, one obtains

$$692 \quad \frac{\partial^2}{\partial \mu^2}(A \cos \phi) = a \frac{\partial}{\partial \mu}(\bar{q}(\mu, t) - Q(\mu, t)), \quad (\text{B1})$$

693 where $Q(\mu, t)$ is absolute vorticity in equivalent latitude [Nakamura and Zhu 2010
 694 Eq. (19)]. From the definition of zonal-mean absolute vorticity

$$695 \quad \frac{\partial^2}{\partial \mu^2}(\bar{u} \cos \phi) = a \left(2\Omega - \frac{\partial \bar{q}}{\partial \mu} \right). \quad (\text{B2})$$

696 Adding (B1) and (B2) then using (3b)

$$697 \quad \frac{\partial^2}{\partial \mu^2}(U_{REF} \cos \phi) = a \left(2\Omega - \frac{\partial Q}{\partial \mu} \right). \quad (\text{B3})$$

698 We first calculate $Q(\mu, t)$ by inverting the area-absolute vorticity relation

$$699 \quad S(Q) = 2\pi a^2(1 - \mu), \quad (\text{B4})$$

700 where $S(Q)$ is the area of domain in which $q \geq Q$. This is evaluated with equally
 701 spaced 1024 bins of Q between Q_{\max} and Q_{\min} , the maximum and minimum values of
 702 absolute vorticity. $S(Q)$ is then inverted for $Q(\mu)$ on equally spaced μ (1024
 703 latitudes), using linear interpolation. Then we calculate the gradient of Q with a
 704 finite difference method and evaluate the right hand side of (B3) on each μ . Finally
 705 we solve for $U_{REF} \cos \phi$ using a tridiagonal solver with the boundary conditions

706 $U_{REF} \cos \phi = 0$ at $\mu = \pm 1$.

707 Once $U_{REF} \cos \phi$ is obtained, we interpolate the result onto Gaussian latitudes

708 on which the model's variables are evaluated, and subtract $\bar{u} \cos \phi$ to obtain $A \cos \phi$.

References

709

710 Andrews, D. G., and M. E. McIntyre, 1976: Planetary waves in horizontal and vertical

711 shear: The generalized Eliassen-Palm relation and the mean zonal

712 acceleration. *J. Atmos. Sci.*, **33**, 2031-2048.

713 ——— and ———, 1978a: An exact theory of nonlinear waves on a Lagrangian-mean

714 flow. *J. Fluid Mech.*, **89**, 609-646.

715 ——— and ———, 1978b: Wave-action and its relatives. *J. Fluid Mech.*, **89**, 647-664.

716 Andrews, D. G., J. R. Holton, and C. B. Leovy, 1987: *Middle Atmosphere Dynamics*.

717 Academic Press, 489pp pp.

718 Boyd, J. P., 1976: The noninteraction of waves with the zonally averaged flow on a

719 spherical earth and the interrelationships of eddy fluxes of energy, heat and

720 momentum. *J. Atmos. Sci.*, **33**, 2285-2291.

721 Bretherton, F. P., and C. J. R. Garrett, 1968: Wavetrains in inhomogeneous moving

722 media. *Proc. Roy. Soc. London Ser. A*, **302**, 529-554.

723 Buchwald, V. T., 1972: Energy and energy flux in planetary waves. *Proc. Roy. Soc.*

724 *London Ser. A*, **328**, 37-48.

725 Bühler, O., 2009: *Wave and mean flows*. Cambridge University Press, 341pp.

726 Charney, J. G., and P. G. Drazin, 1961: Propagation of planetary-scale disturbances

727 from the lower into the upper atmosphere. *J. Geophys. Res.*, **66**, 83-110.

728 Dickinson, R. E., 1969: Theory of planetary wave-zonal flow interaction. *J. Atmos. Sci.*,

729 **26**, 73-81.

730 Edmon, H. J., B. J. Hoskins, and M. E. McIntyre, 1980: Eliassen-Palm cross sections for

731 the troposphere. *J. Atmos. Sci.*, **37**, 2600-2616.

732 Grimshaw, R., 1984: Wave action and wave-mean flow interaction, with application
733 to stratified shear flows. *Ann. Rev. Fluid Mech.*, **16**, 11-44.

734 Haurwitz, B., 1940: The motion of atmospheric disturbances on the spherical earth. *J.*
735 *Mar. Res.*, **3**, 254-267.

736 Hayes, W. D., 1970: Conservation of action and modal wave action. *Proc. Roy. Soc.*
737 *London Ser. A*, **320**, 187-208.

738 Haynes, P. H., 1988: Forced, dissipative generalizations of finite-amplitude wave-
739 activity conservation relations for zonal and nonzonal basic flows. *J. Atmos.*
740 *Sci.*, **45**, 2352-2362.

741 Haynes, P. H., and M. E. McIntyre, 1987: On the representation of rossby-wave critical
742 layers and wave breaking in zonally truncated models. *J. Atmos. Sci.*, **44**, 2359-
743 2382.

744 Held, I. M., 1983: Stationary and quasi-stationary eddies in the extratropical
745 troposphere: theory. *Large-scale dynamical processes in the atmopshere*, B. J.
746 Hoskins, and R. P. Pearce, Eds., Academic Press, 127-167.

747 —, 1985: Pseudomomentum and the orthogonality of modes in shear flows. *J.*
748 *Atmos. Sci.*, **42**, 2280-2288.

749 Held, I. M., and P. J. Phillips, 1987: Linear and nonlinear barotropic decay on the
750 sphere. *J. Atmos. Sci.*, **44**, 200-207.

751 Hoskins, B. J., and D. J. Karoly, 1981: The steady linear response of a spherical
752 atmosphere to thermal and orographic forcing. *J. Atmos. Sci.*, **38**, 1179-1196.

753 Killworth, P. D., and M. E. McIntyre, 1985: Do Rossby-wave critical layers absorb,
754 reflect, or over-reflect? *J. Fluid Mech.*, **161**, 449-492.

755 Longuet_Higgins, M. S., 1964: Planetary waves on a rotating sphere. *Proc. Roy. Soc.*
756 *London Ser. A*, **279**, 446-473.

757 ———, 1965: Planetary waves on a rotating sphere. II. *Proc. Roy. Soc. London Ser. A*,
758 **284**, 40-68.

759 Matsuno, T., 1971: A dynamical model of the stratospheric sudden warming. *J. Atmos.*
760 *Sci.*, **28**, 1479-1494.

761 McIntyre, M. E., and T. N. Palmer, 1983: Breaking planetary waves in the stratosphere.
762 *Nature*, **305**, 593-600.

763 McIntyre, M. E., and T. G. Shepherd, 1987: An exact local conservation theorem for
764 finite-amplitude disturbances to nonparallel shear flows, with remarks on
765 Hamiltonian-structure and on Arnold stability theorems. *J. Fluid Mech.*, **181**,
766 527-565.

767 Nakamura, N., and A. Solomon, 2010: Finite-amplitude wave activity and mean flow
768 adjustments in the atmospheric general circulation. Part I: Quasigeostrophic
769 theory and analysis. *J. Atmos. Sci.*, **67**, 3967-3983.

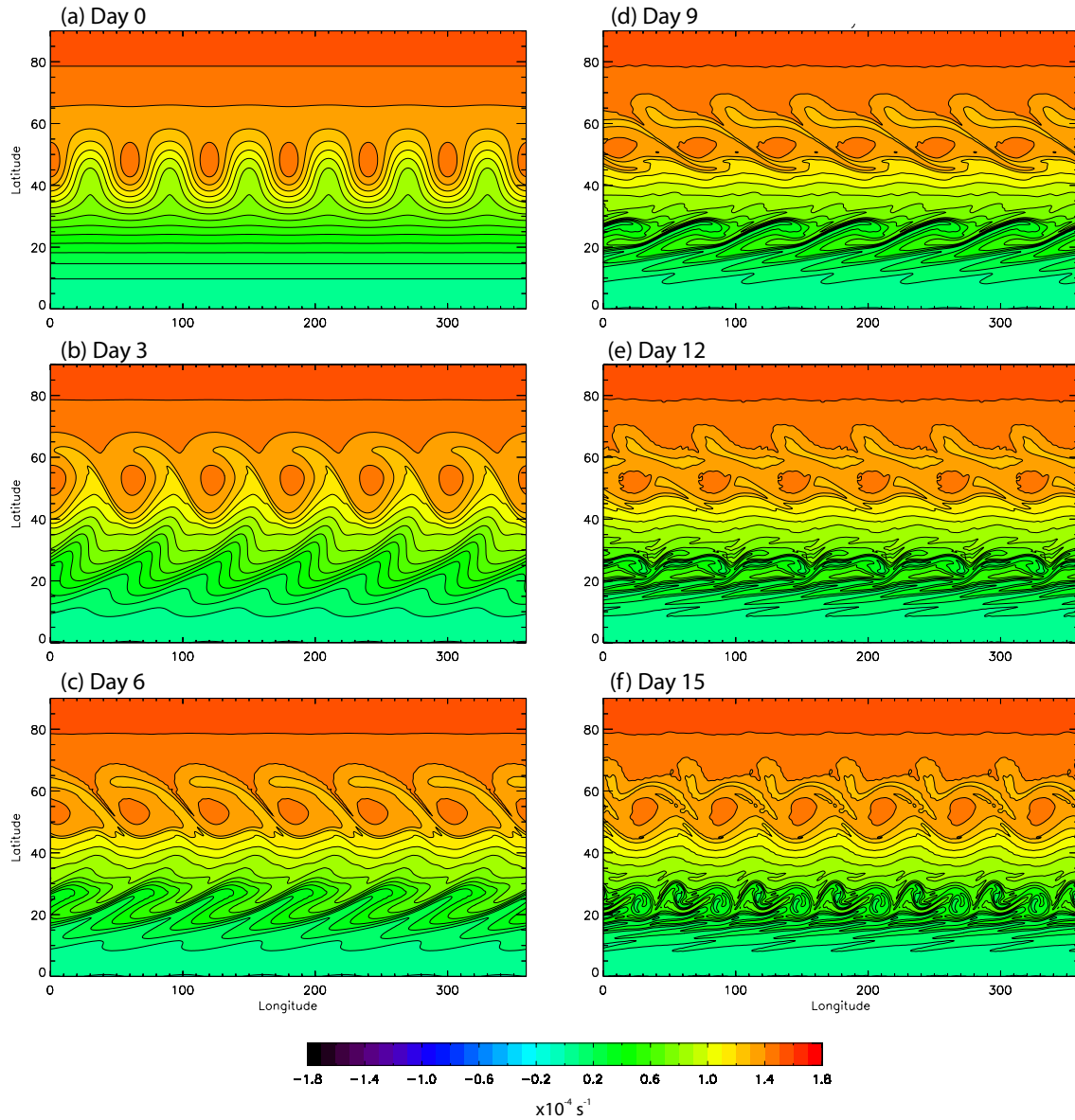
770 Nakamura, N., and D. Zhu, 2010: Finite-amplitude wave activity and diffusive flux of
771 potential vorticity in eddy-mean flow interaction. *J. Atmos. Sci.*, **67**, 2701-2716.

772 Platzman, G. W., 1968: The Rossby wave. *Quart. J. Roy. Meteor. Soc.*, **94**, 225-248.

773 Randel, W. J., and I. M. Held, 1991: Phase speed spectra of transient eddy fluxes and
774 critical layer absorption. *J. Atmos. Sci.*, **48**, 688-697.

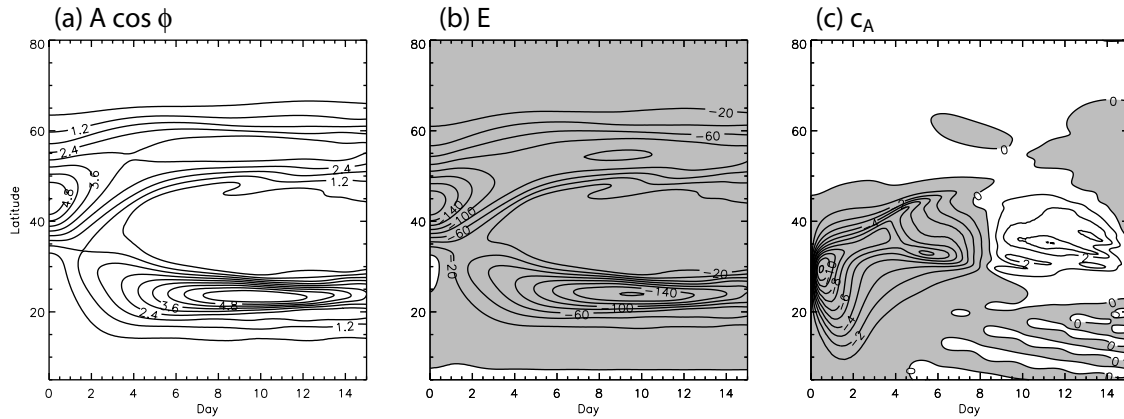
775 Rossby, C. G., 1939: Relation between variations in the intensity of the zonal
776 circulation of the atmosphere and the displacements of the semipermanent
777 centers of action. *J. Mar. Res.*, **2**, 38-55.

- 778 Salmon, R., 2013: An alternative view of generalized Lagrangian mean theory. *J. Fluid*
779 *Mech.*, **719**, 165-182.
- 780 Seliger, R. L., and G. B. Whitham, 1968: Variational principles in continuum
781 mechanics. *Proc. Roy. Soc. London Ser. A*, **305**, 1-25.
- 782 Solomon, A., and N. Nakamura, 2012: An exact Lagrangian-mean wave activity for
783 finite-amplitude disturbances to barotropic flow on the sphere. *J. Fluid Mech.*,
784 **693**, 69-92.
- 785 Sturrock, P. A., 1961: Energy-momentum tensor for plane waves. *Phys. Rev.*, **121**, 18-
786 19.
- 787 Whitham, G. B., 1965: A general approach to linear and non-linear dispersive waves
788 using a Lagrangian. *J. Fluid Mech.*, **22**, 273-283.
- 789 —, 1970: Two-timing, variational principles and waves. *J. Fluid Mech.*, **44**, 373-395.
- 790 —, 1974: *Linear and nonlinear waves*. New York: Wiley-Interscience. 656pp.



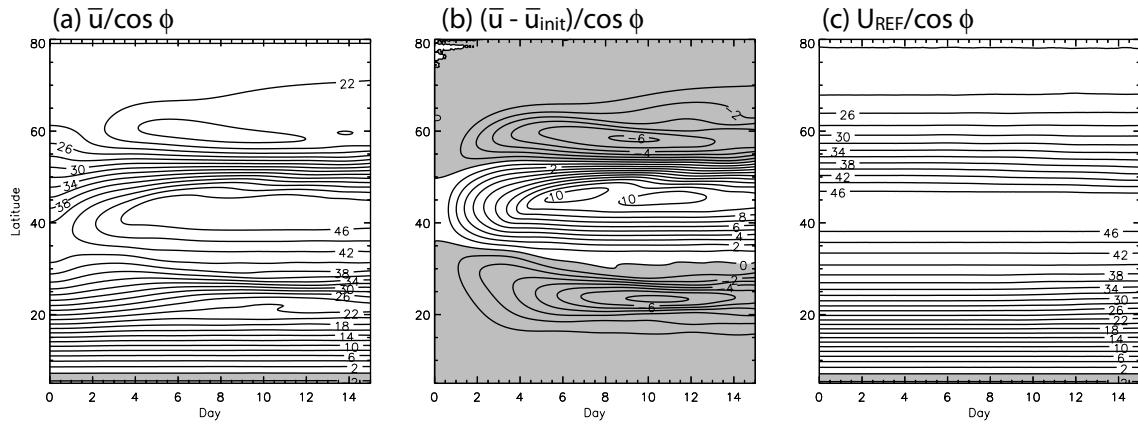
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792 **Figure 1** Evolution of absolute vorticity in the Northern Hemisphere during
 793 nonlinear barotropic decay described in section 3. Abscissa is longitude and ordinate
 794 is latitude. (a) initial condition, (b) day 3, (c) day 6, (d) day 9, (e) day 12, (f) day 15.



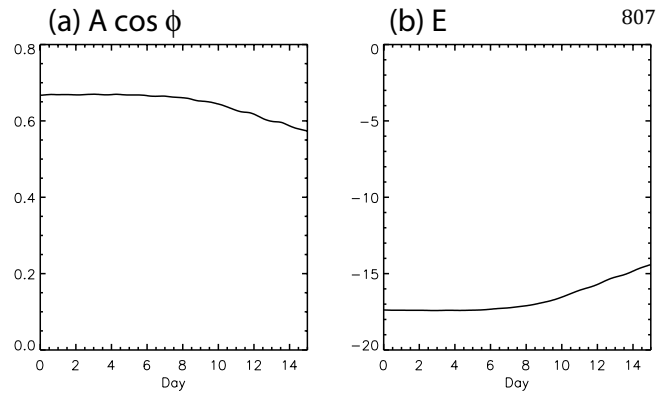
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796 **Figure 2** Wave properties as functions of time and latitude during the same
 797 numerical simulation as in Fig.1. (a) Angular pseudomomentum density $A \cos \phi$,
 798 contour interval = 0.6 m s^{-1} (b) pseudoenergy density E , contour interval = $20 \text{ m}^2 \text{ s}^{-2}$
 799 (c) effective transport velocity c_A (= group velocity in small-amplitude limit),
 800 contour interval = 1 m s^{-1} . In (b) and (c), negative values are shaded. See text for the
 801 definitions of variables.



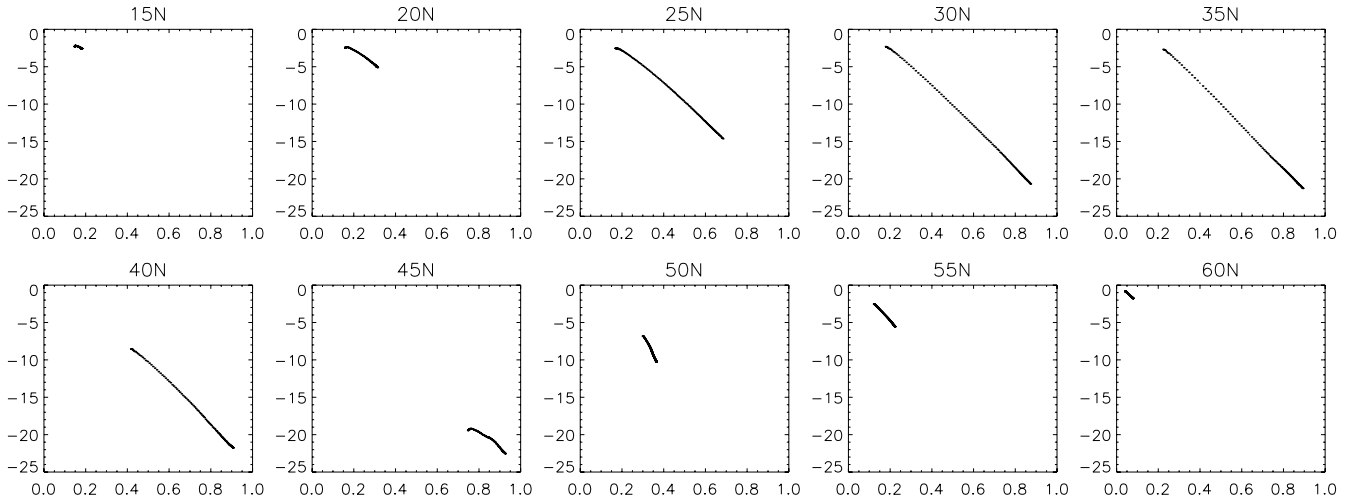
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803 **Figure 3** Same as Fig.2 but for the mean flow properties. (a) Zonal-mean angular
 804 velocity $\bar{u} / \cos\phi$, contour interval = 2 ms^{-1} (b) $\bar{u} / \cos\phi$ minus its initial value,
 805 contour interval = 1 ms^{-1} (c) Angular velocity of the reference state flow $U_{REF} / \cos\phi$,
 806 contour interval = 2 ms^{-1} . Negative values are shaded.



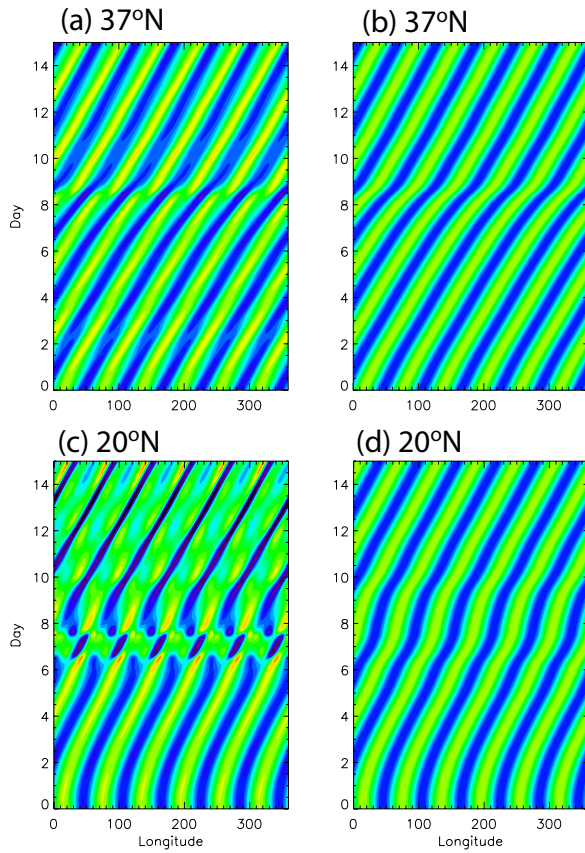
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809 **Figure 4** (a) Domain-averaged $A \cos \phi$ as a function of time during the above
 810 simulation. Unit: ms^{-1} . (b) Same as (a) but for domain-averaged E . Unit: m^2s^{-2} .



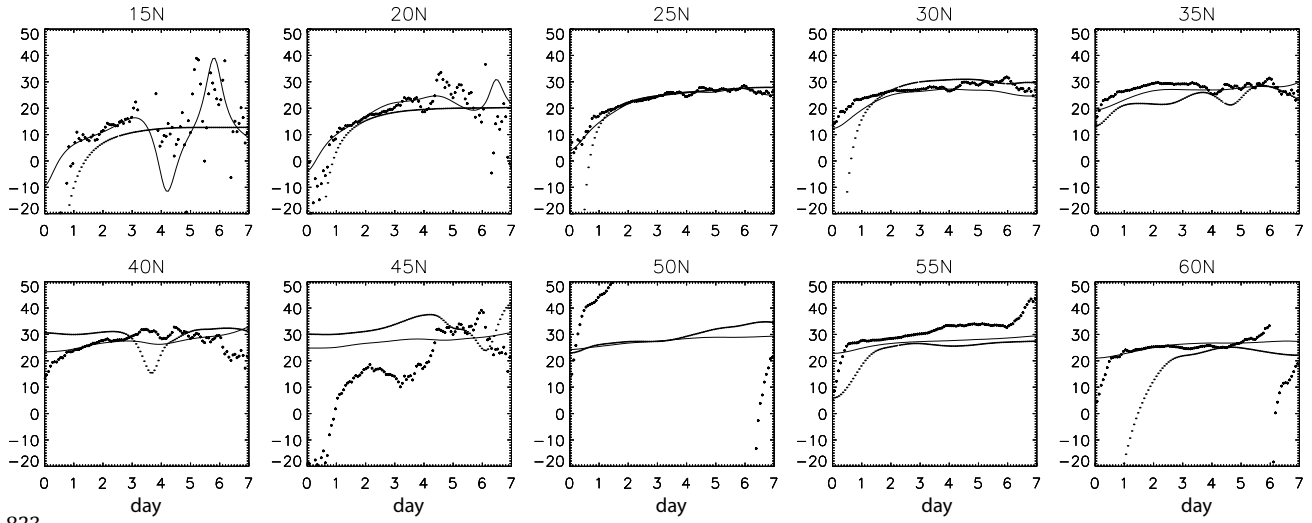
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812 **Figure 5** Scatter plots of A^\dagger (abscissa, ms^{-1}) versus E^\dagger (ordinate, m^2s^{-2}) at 10
 813 different latitudes, each constructed from the first 7 days of simulation. Each frame
 814 show 101 data points, sampled every 6000 s. The wave amplitude is small in the
 815 upper left and large in the lower right. For 50° , 55° , and 60° N, A^\dagger and E^\dagger were
 816 evaluated by southward integration from the North Pole. See Eq. (43) and text for
 817 details.



818

819 **Figure 6** Left: longitude-time plots (Hovmöller diagrams) of the rms-normalized v
 820 velocity for the simulation, at 37° N (top) and 20° N (bottom). Right: Reconstruction
 821 of phase by integrating the ‘observed’ phase speed calculated from Eq. (45). See text
 822 for details.



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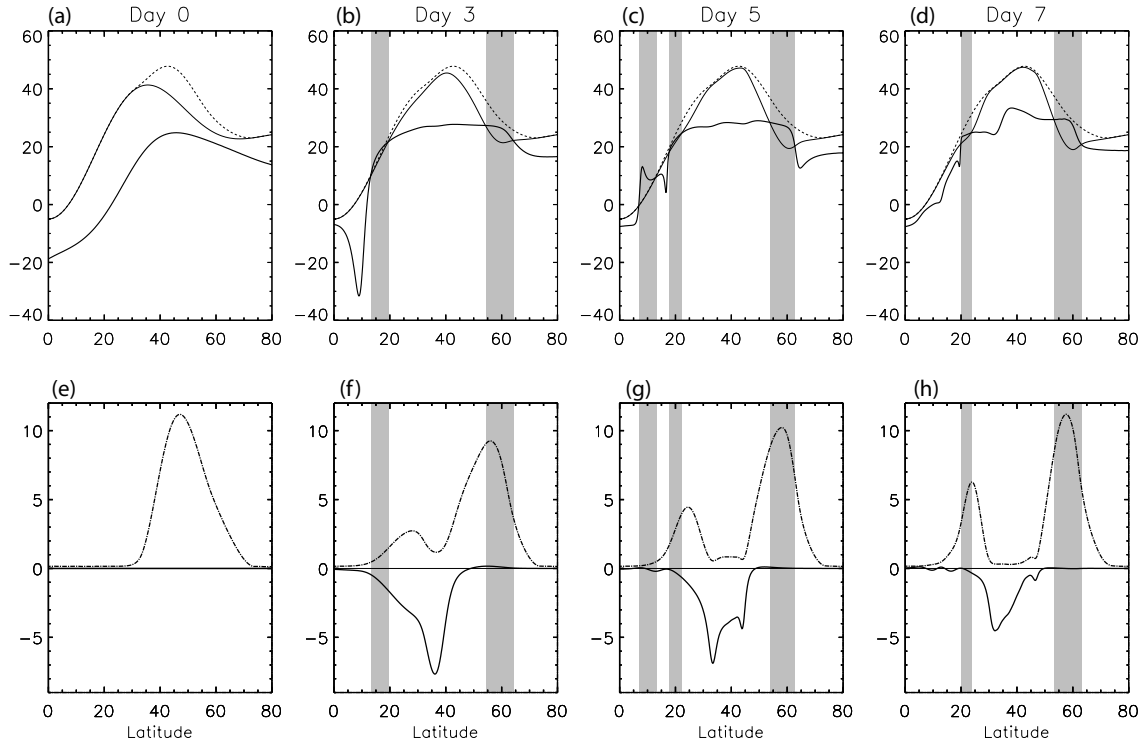
824 **Figure 7** Phase speed of the Rossby wave c during the first 7 days of simulation at

825 10 different latitudes. Thin solid curve: 'observed' phase speed based on Eq. (45).

826 Thick dots: theoretical estimates based on (43) and the $A^\dagger - E^\dagger$ relations in Fig.5.

827 Thin dots: theoretical estimates based on linear theory [(30) and (29)]. The unit of

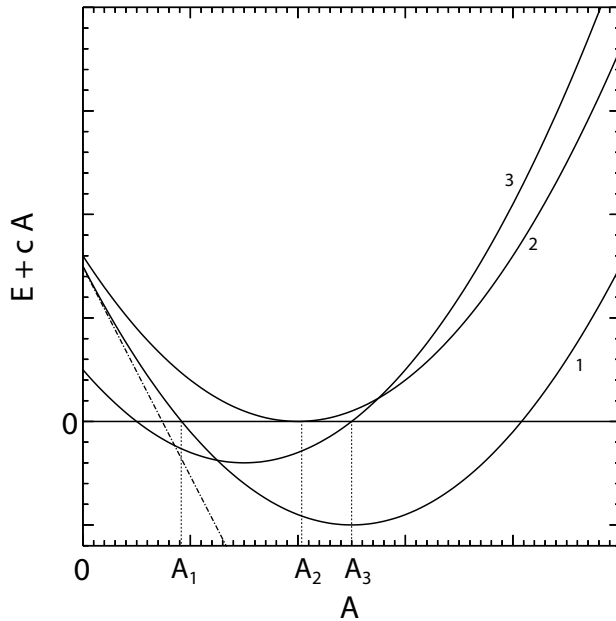
828 the ordinate is ms^{-1} . See text for details.



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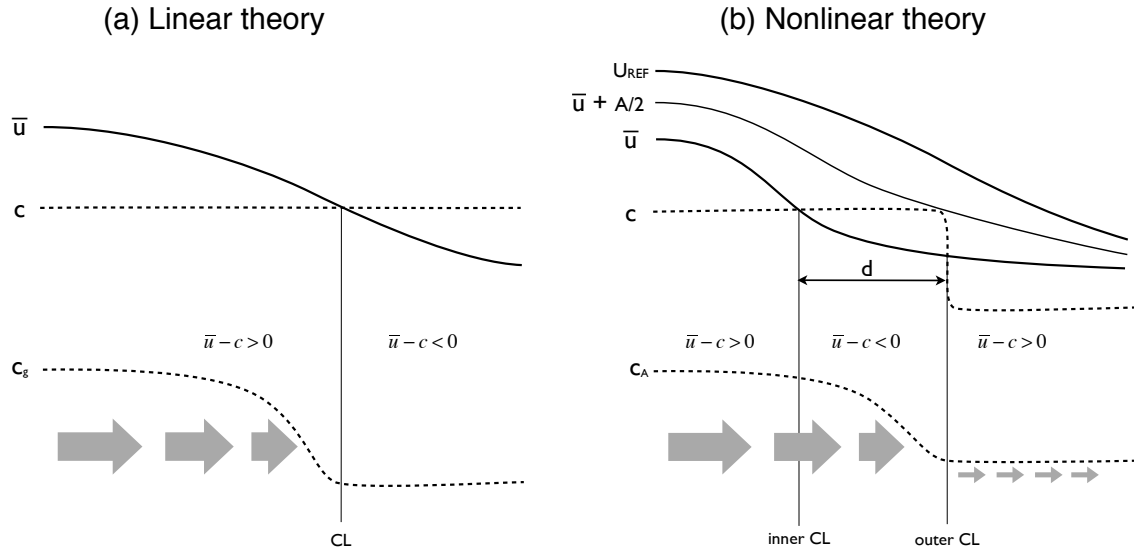
830 **Figure 8** Top row: c_{obs} based on (45) (thick solid curve) and $\bar{u}/\cos\phi$ (thin solid
 831 curve) and $U_{REF}/\cos\phi$ (dashed curve) as functions of latitude at four different
 832 instants during the numerical simulation of nonlinear barotropic decay. (a) Day 0
 833 (b) Day 3 (c) Day 5 (d) Day 7. In the shaded region $c_{obs} \geq \bar{u}/\cos\phi$: the boundaries of
 834 these regions mark the critical lines. Bottom row: same as the top row but for c_A
 835 based on (27) (thick solid curve) and $A/\cos\phi$ (dot-dashed curve). Zero line is also
 836 added. (e) Day 0 (f) Day 3 (g) Day 5 (h) Day 7. The unit of the ordinate is ms^{-1} in
 837 all plots.

838



839

840 **Figure 9** $E + cA$ as quadratic functions of A . Curves 1-3 correspond to three
 841 different latitudes, y_1, y_2, y_3 . The dash-dotted line indicates linear dispersion
 842 relation at y_1 : it is tangent to curve 1 at $A = 0$. The intercept of each curve with the
 843 ordinate equals \bar{e} at that latitude and the horizontal coordinate of the vertex of each
 844 curve is $U_{REF} - c$. See text for details.



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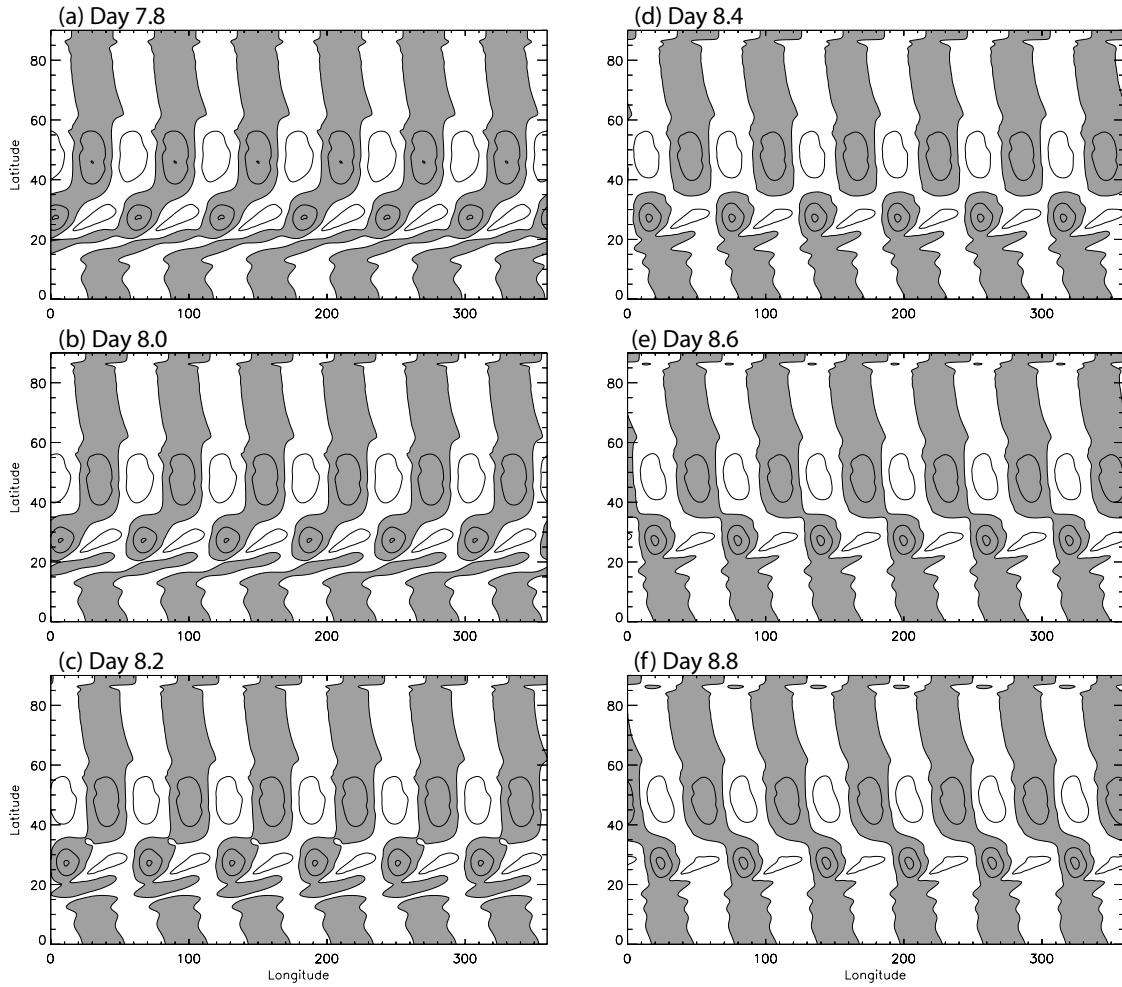
846 **Figure 10** Schematics of mature critical line geometry (horizontal axis is latitude).

847 Critical lines are defined by the intersections of the $\bar{u}(y)$ and $c(y)$ curves. (a)

848 Classical linear theory. Phase speed c is constant across latitude whereas the
 849 meridional group velocity (indicated by gray arrows) drops to zero at the critical line.

850 (b) Nonlinear theory. The inner critical line is nonsingular and transmits the wave
 851 to the right. c is constant up to $c \approx \bar{u} + A/2$, where it drops abruptly. c_A also nearly
 852 vanishes here. The outer critical line coincides with the latitude of sharp falloff of c .

853 The distance between the two critical lines is denoted by d . See text for details.



854

855 **Figure 11** Reconfiguration of phase lines in v -velocity from day 7.8 to day 8.8.

856 Notice the breakup and reconnection of phase lines around $15^\circ - 20^\circ$ N and 35° N.

857 Contour interval is 1 ms^{-1} and negative values are shaded.